



# Semillero de Investigación “Hands - on” Computer Vision





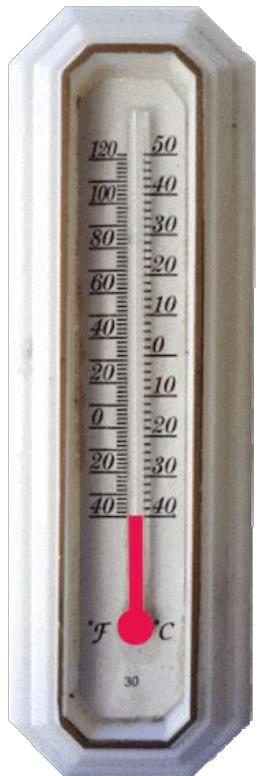
# SESIÓN 8: IMÁGENES TÉRMICAS

# Contenidos

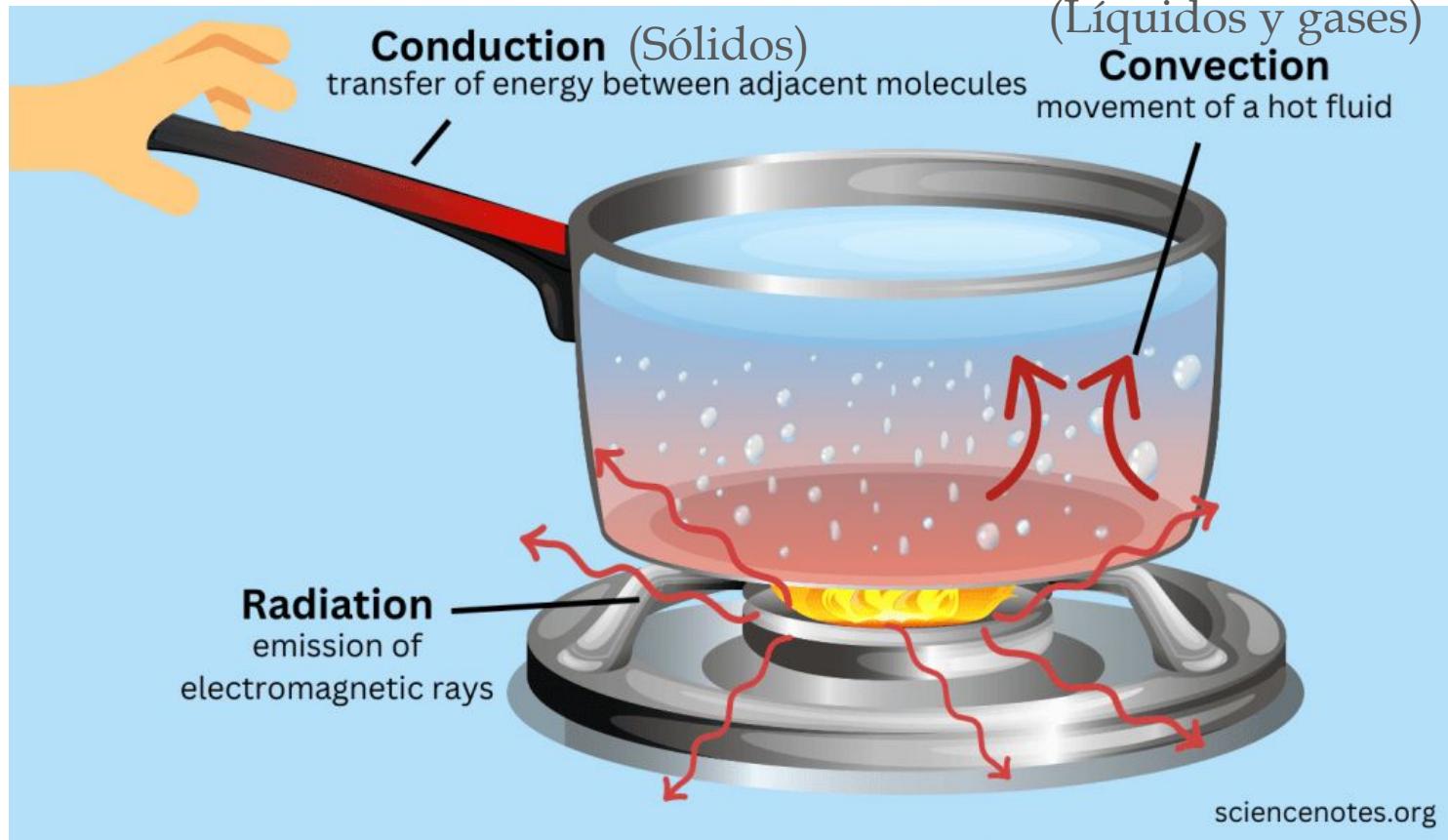
1. Imágenes térmicas y aplicaciones
2. Luz y radiación
  - a. Ley de Kirchhoff
  - b. Radiación de cuerpo negro (Planck)
  - c. Emisividad y atenuación
3. Imágenes espectrales térmicas
4. Thermal & CV
5. Hands-on Thermal Imaging



# ¿Cómo podemos medir la temperatura?



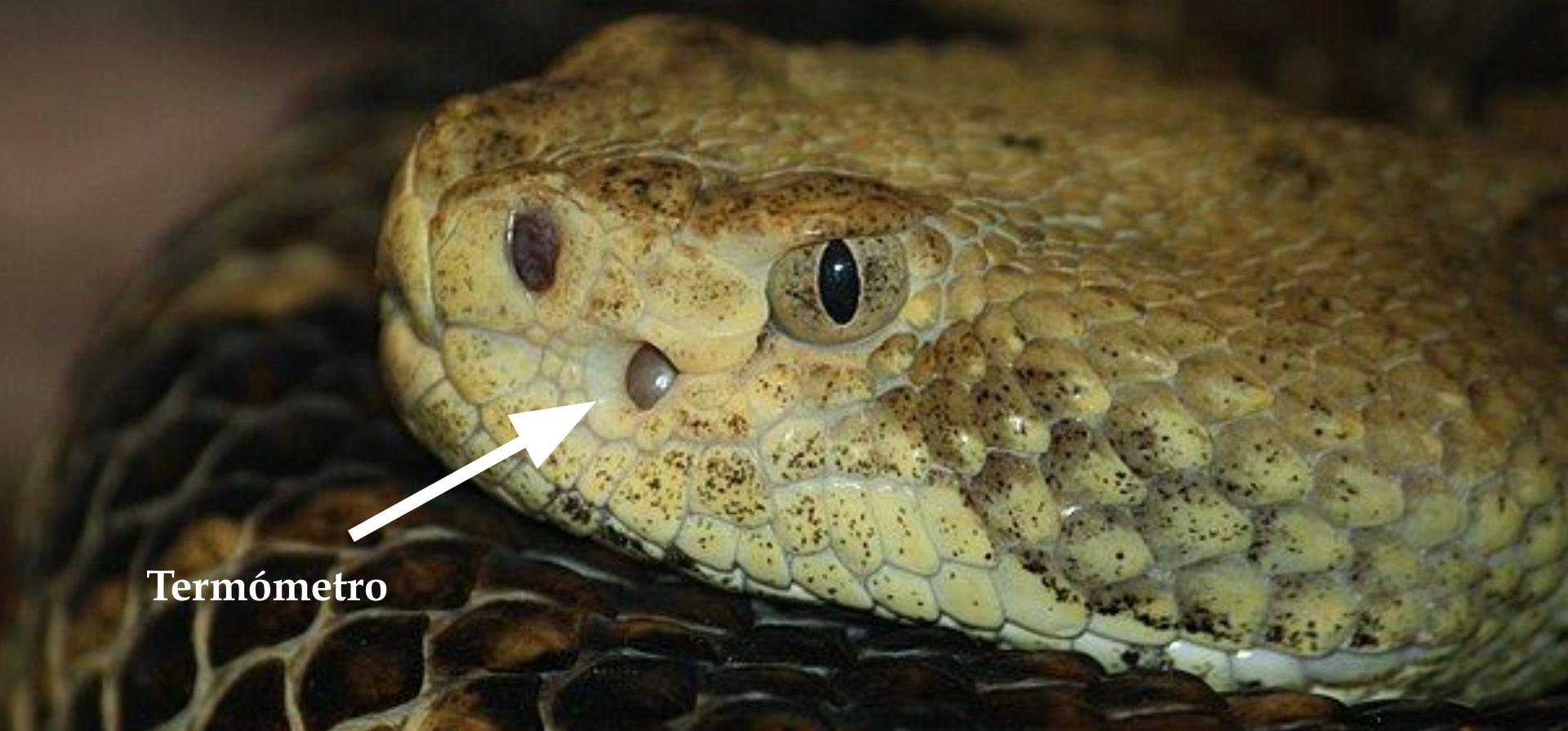
# Transferencia de calor



La conducción requiere contacto. La convección requiere flujo de fluido. La radiación no requiere ningún medio.

# Sensores térmicos en la naturaleza

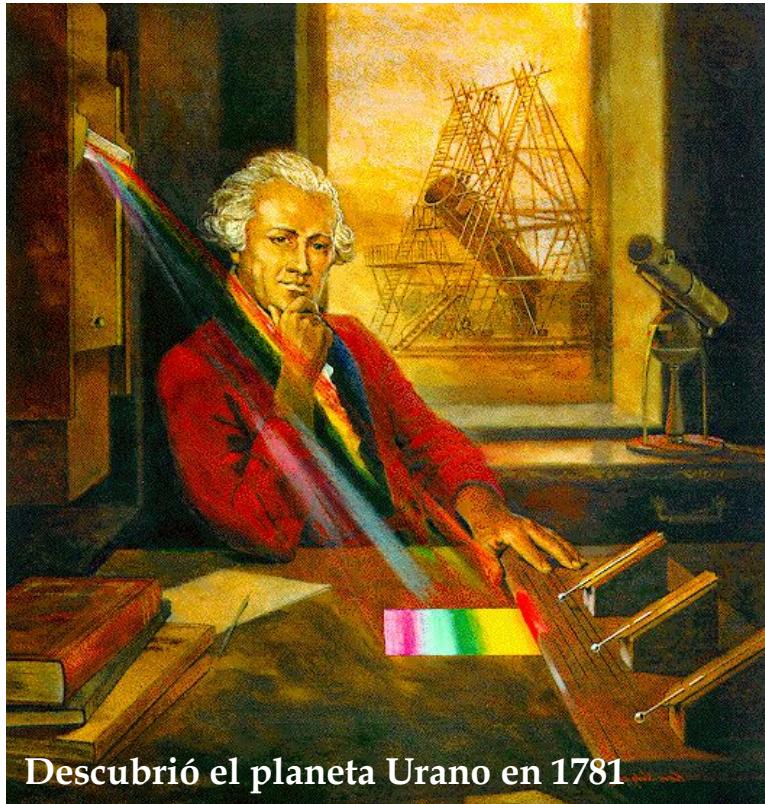
Víboras: pitones, boas y serpientes de cascabel



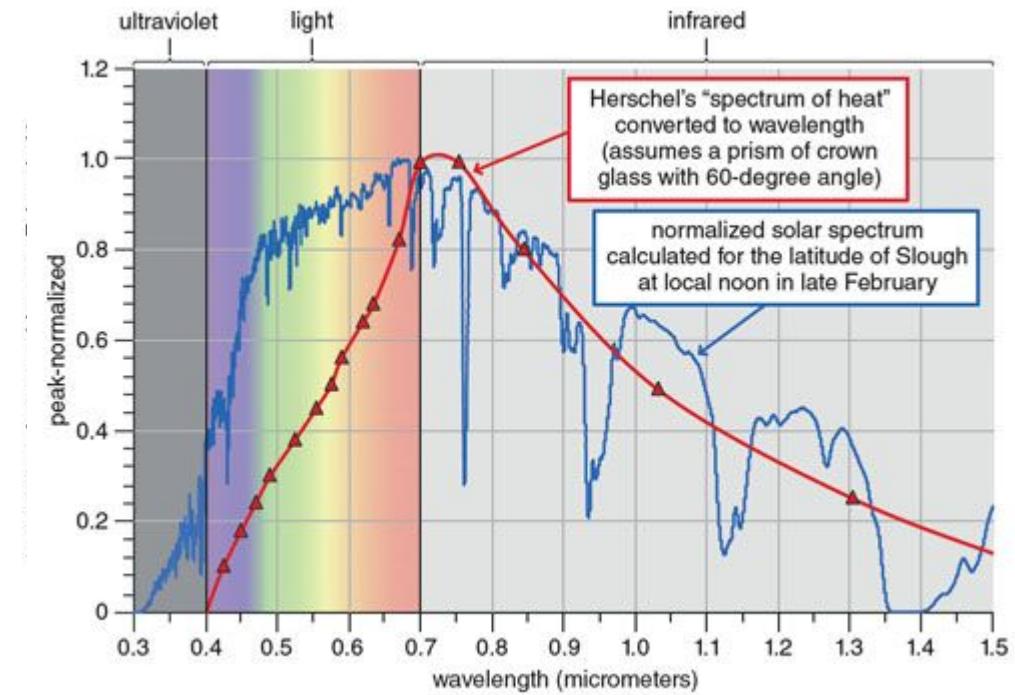
Termómetro

# 1. Imágenes Térmicas

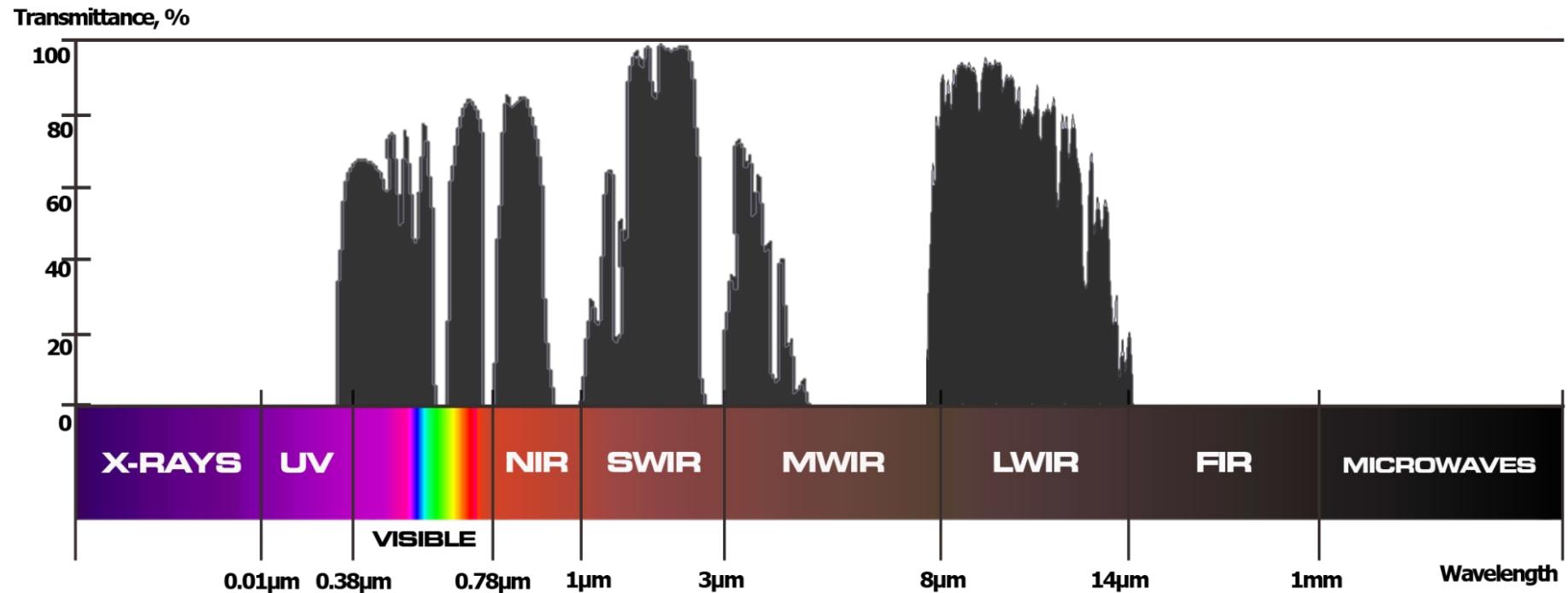
# Infrared light discovery (Frederick W Herschel in 1800)



Descubrió el planeta Urano en 1781

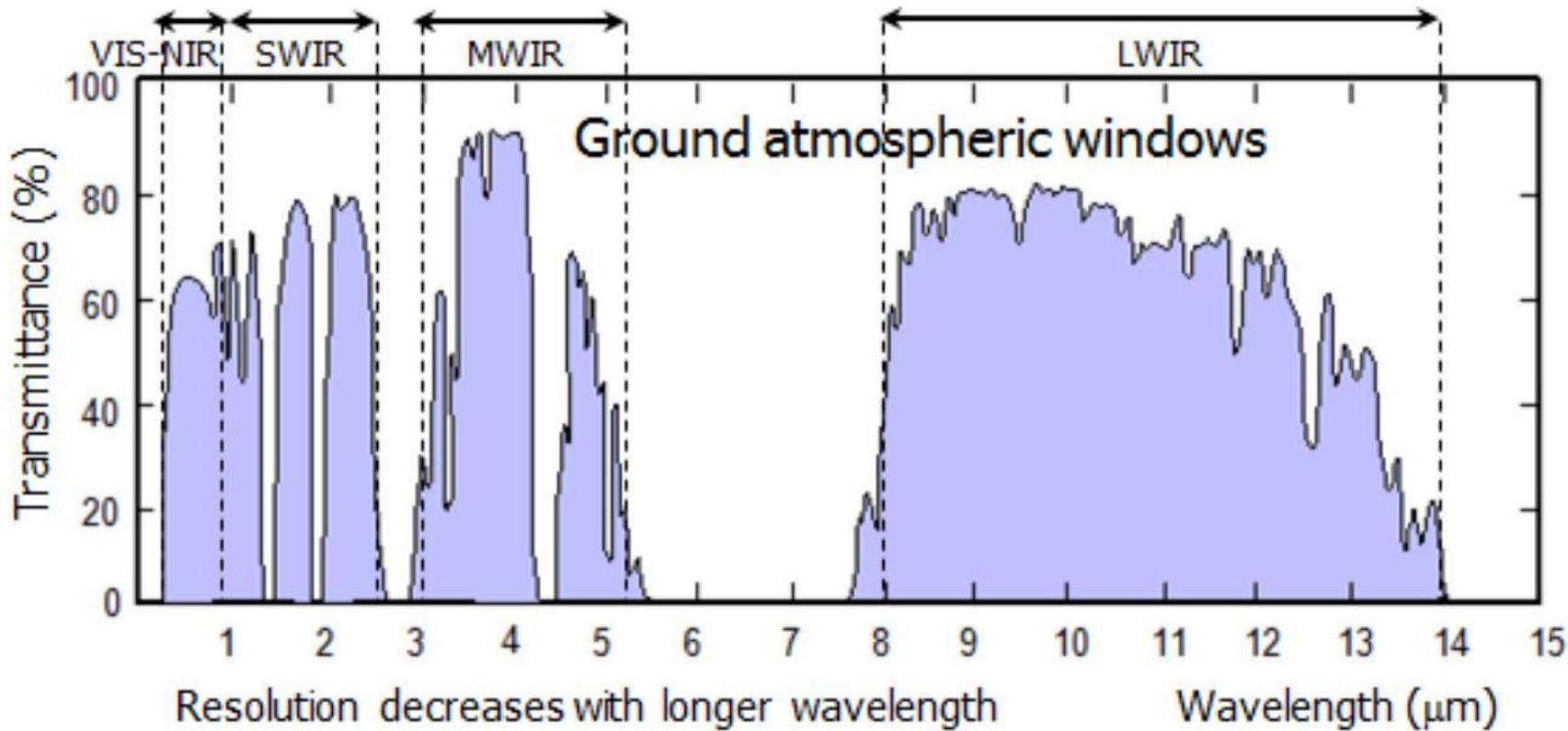


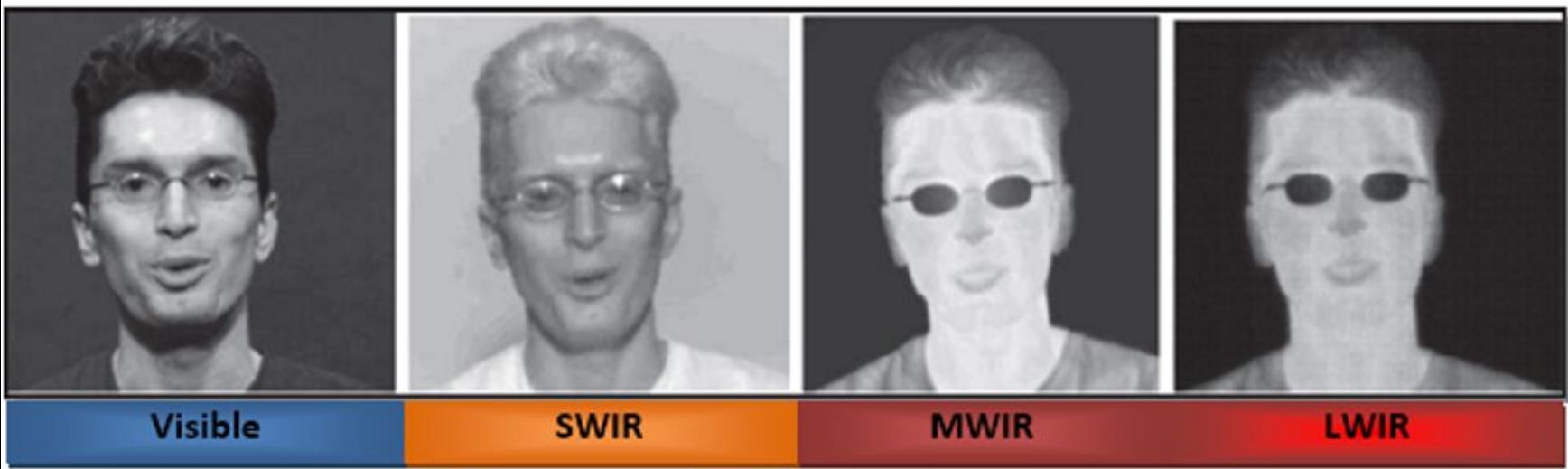
# Espectro electromagnético



Térmico = LWIR

# Importancia de la Atmósfera





**Visible**

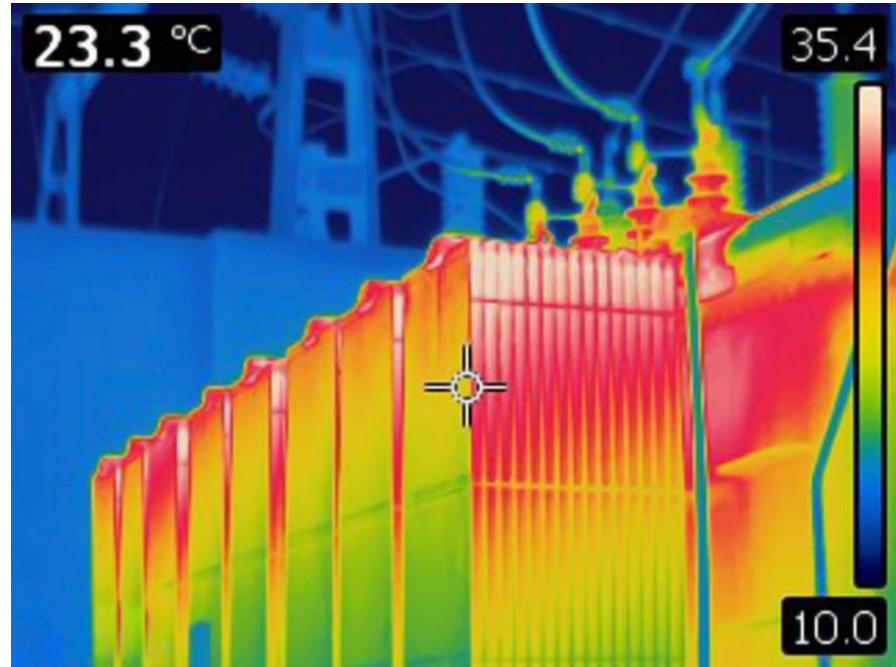
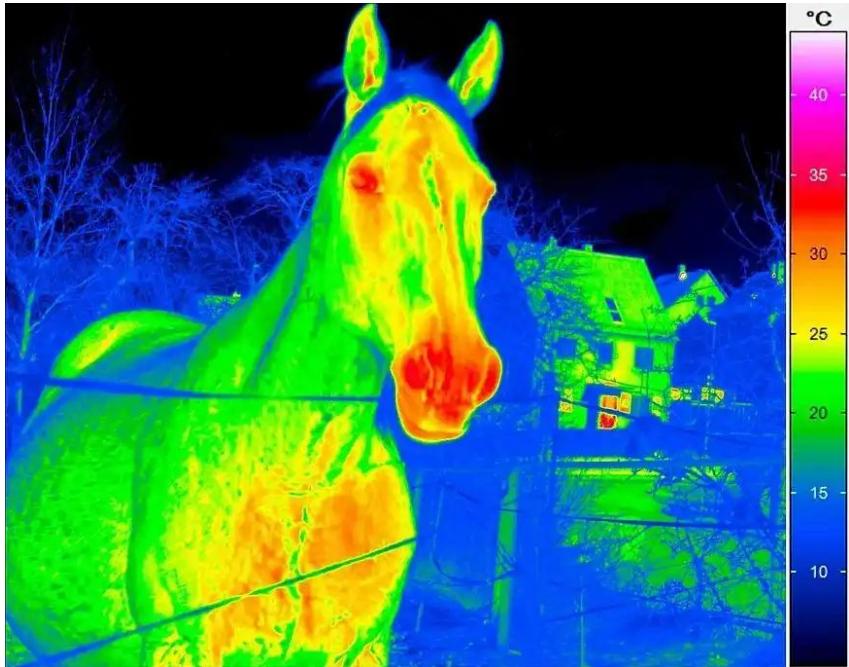
**SWIR**

**MWIR**

**LWIR**

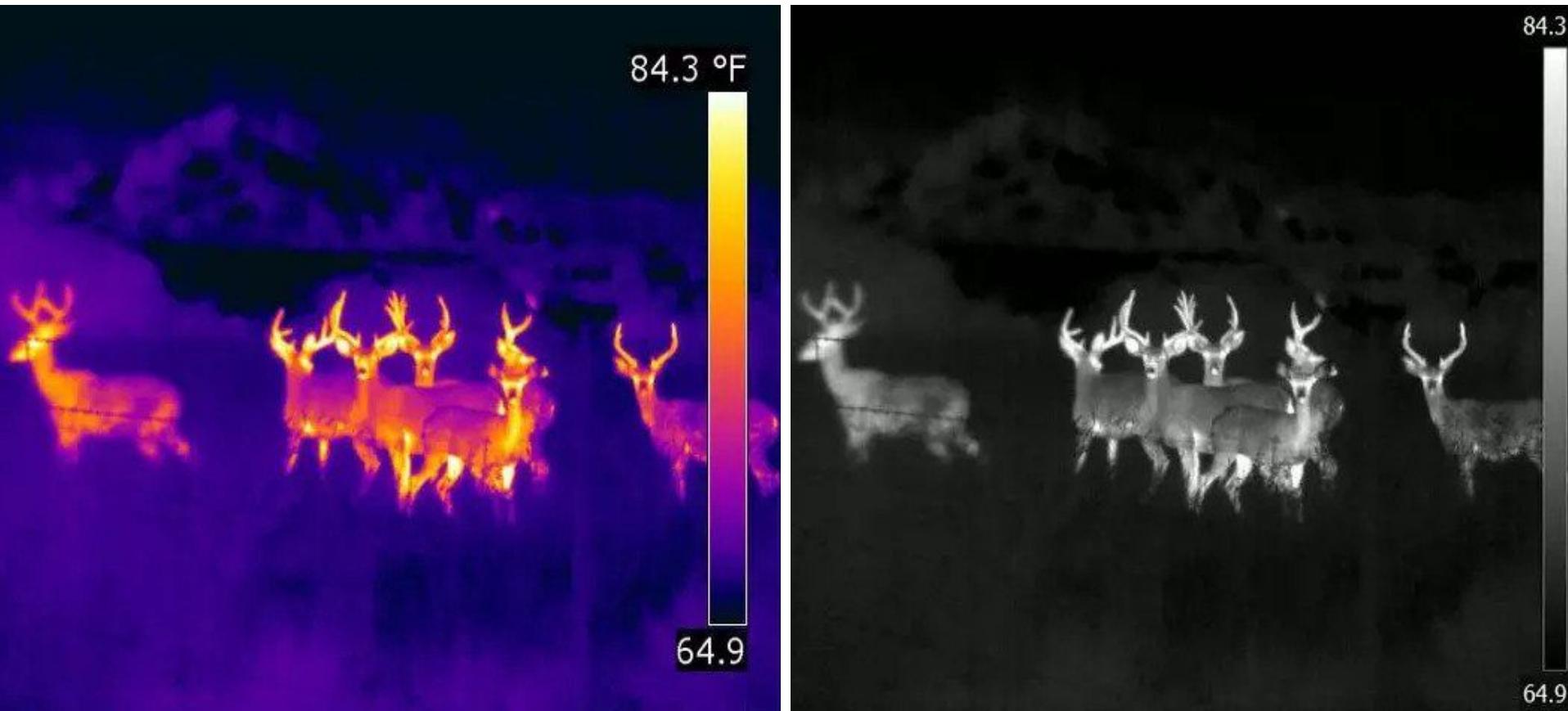


# Imágenes térmicas



- Un termograma es una representación digital de la radiación térmica emitida por los objetos
- Las cámaras termográficas permiten detectar esta radiación en forma de luz infrarroja
- Una imagen térmica nos permite detectar de forma remota la temperatura de un objeto

# Imágenes térmicas

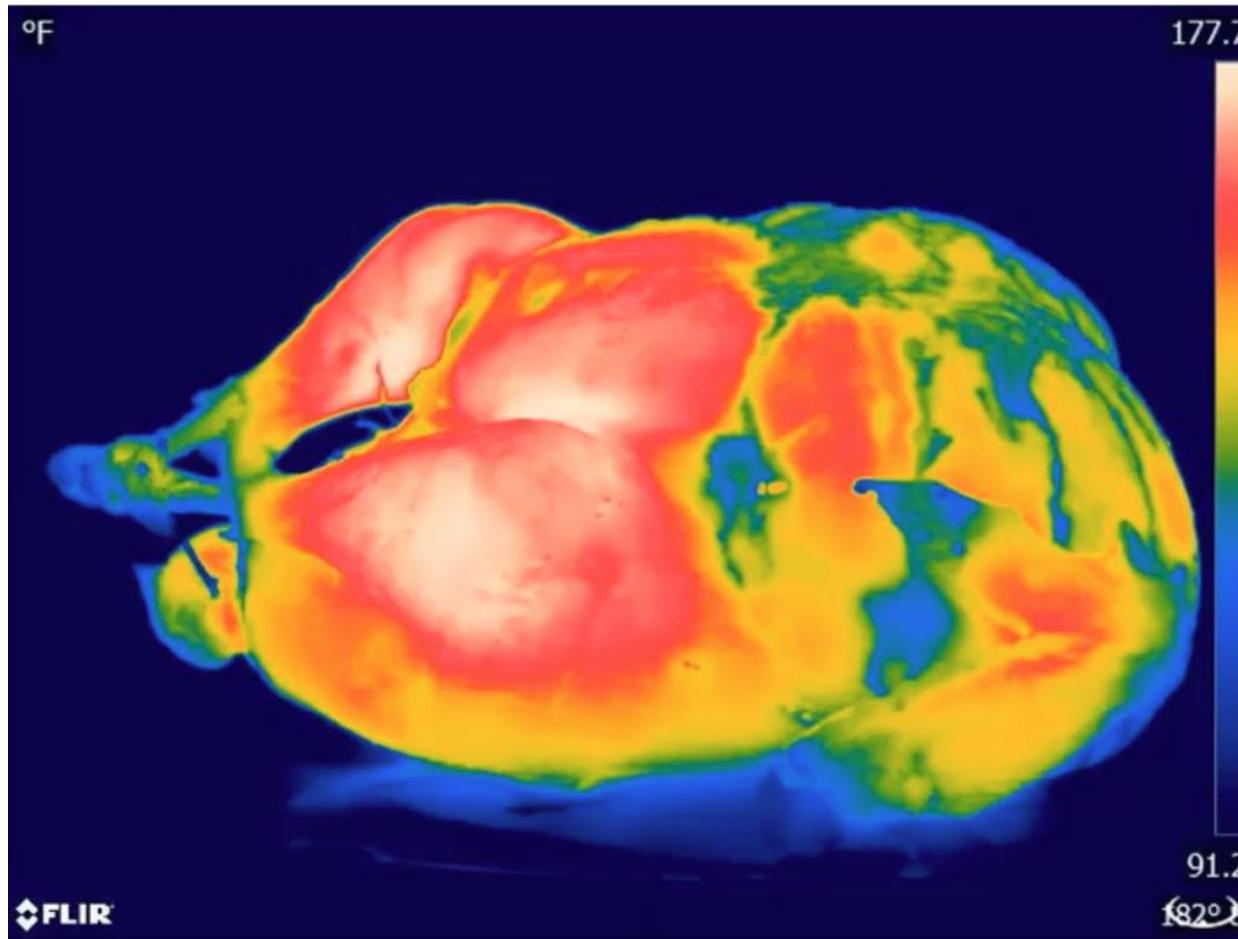


**FLIR** Pioneer





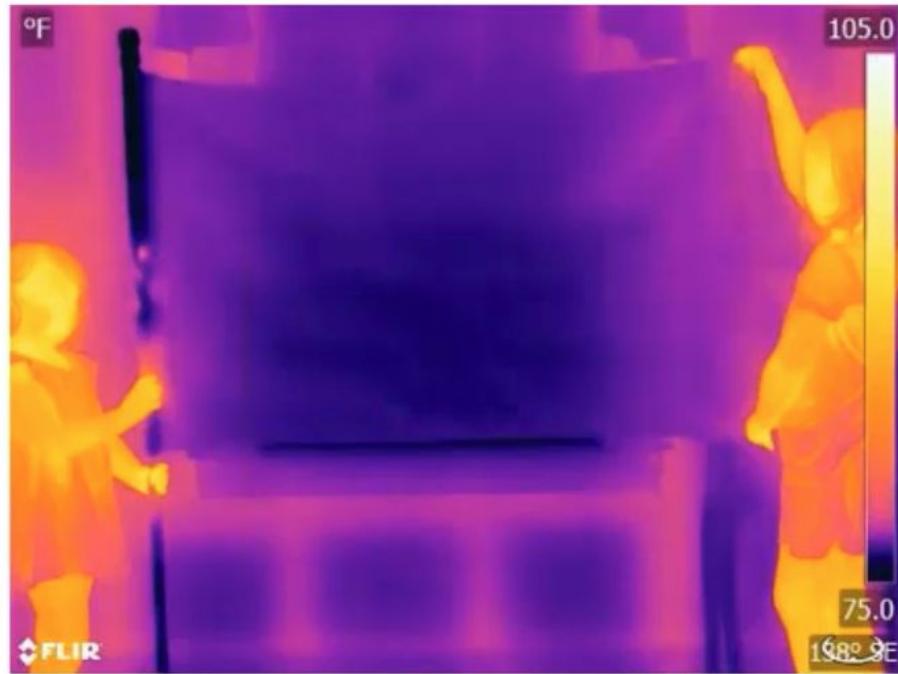
Las alas se enfriaron más rápido (heat capacity, humedad)



# Por qué la bandera se ve diferente?



Afuera de casa

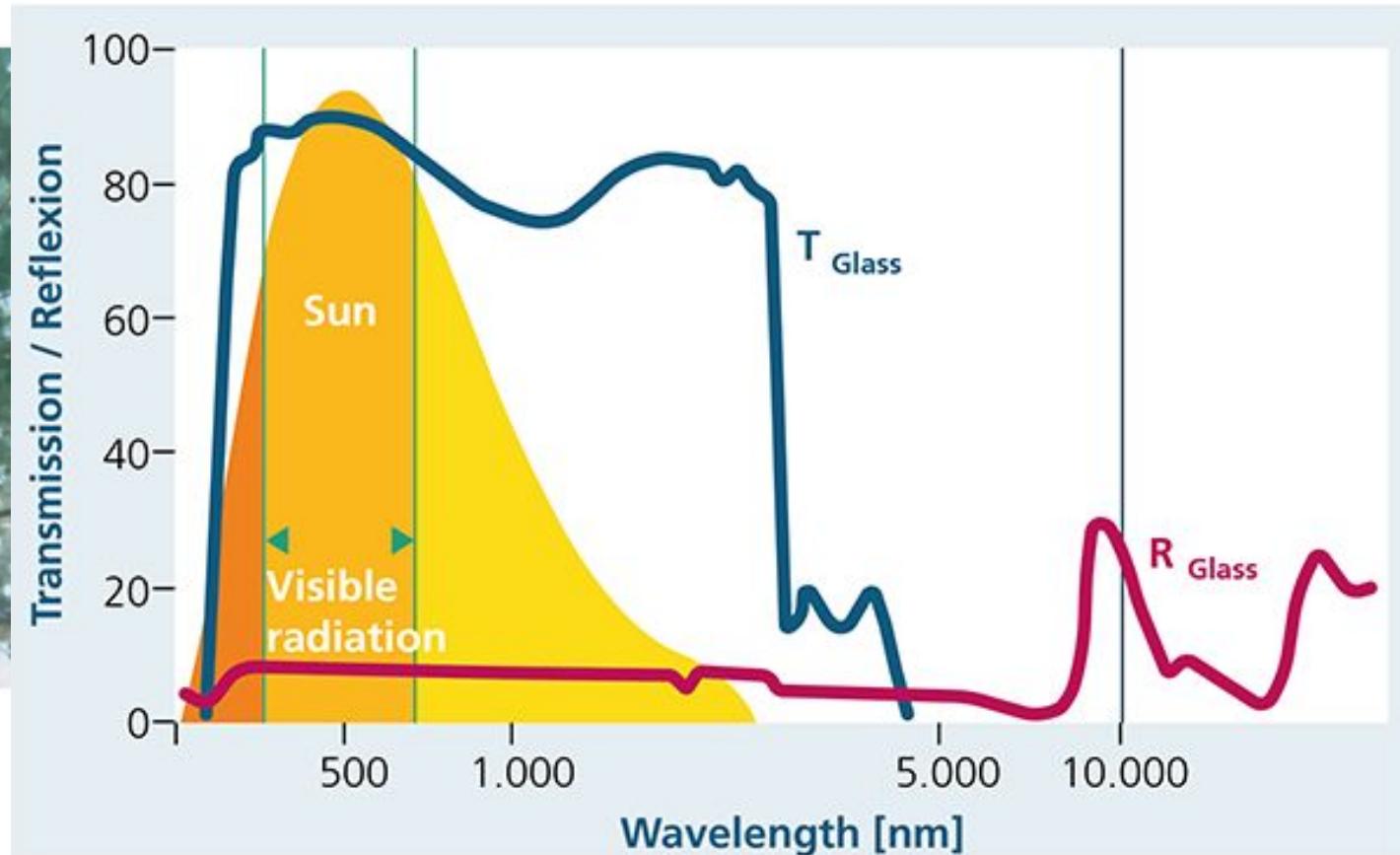
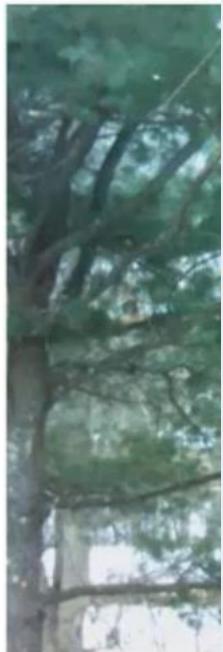


Dentro de casa

# Metales reflejan luz y pueden ser difíciles de medir



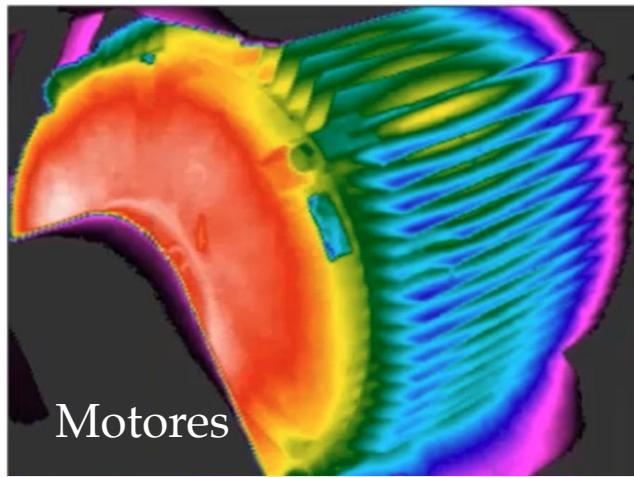
# Los vidrios son opacos en el rango térmico



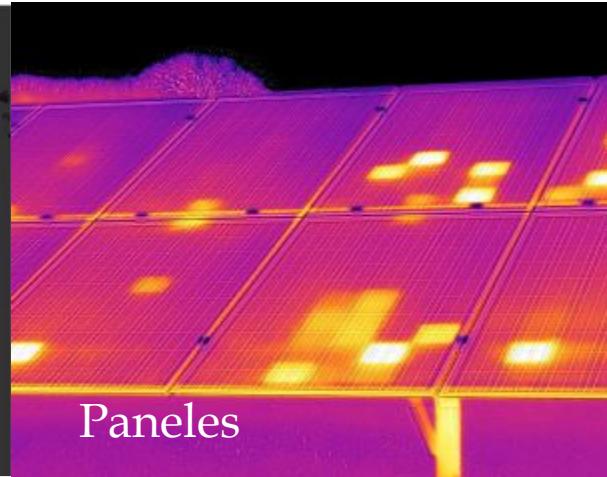
# Aplicaciones



Tanques



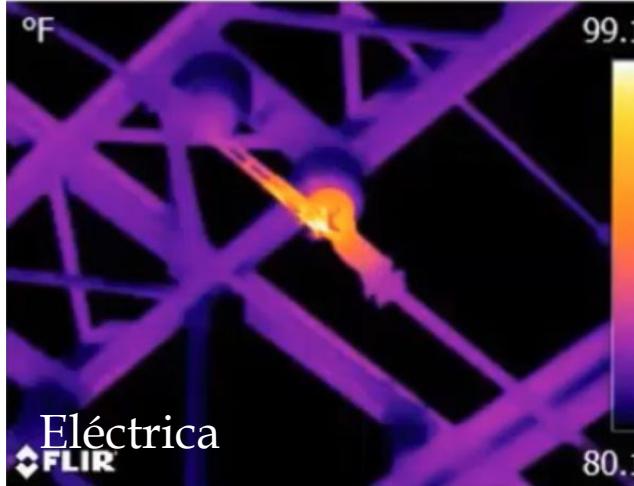
Motores



Paneles



Eléctrica



Eléctrica

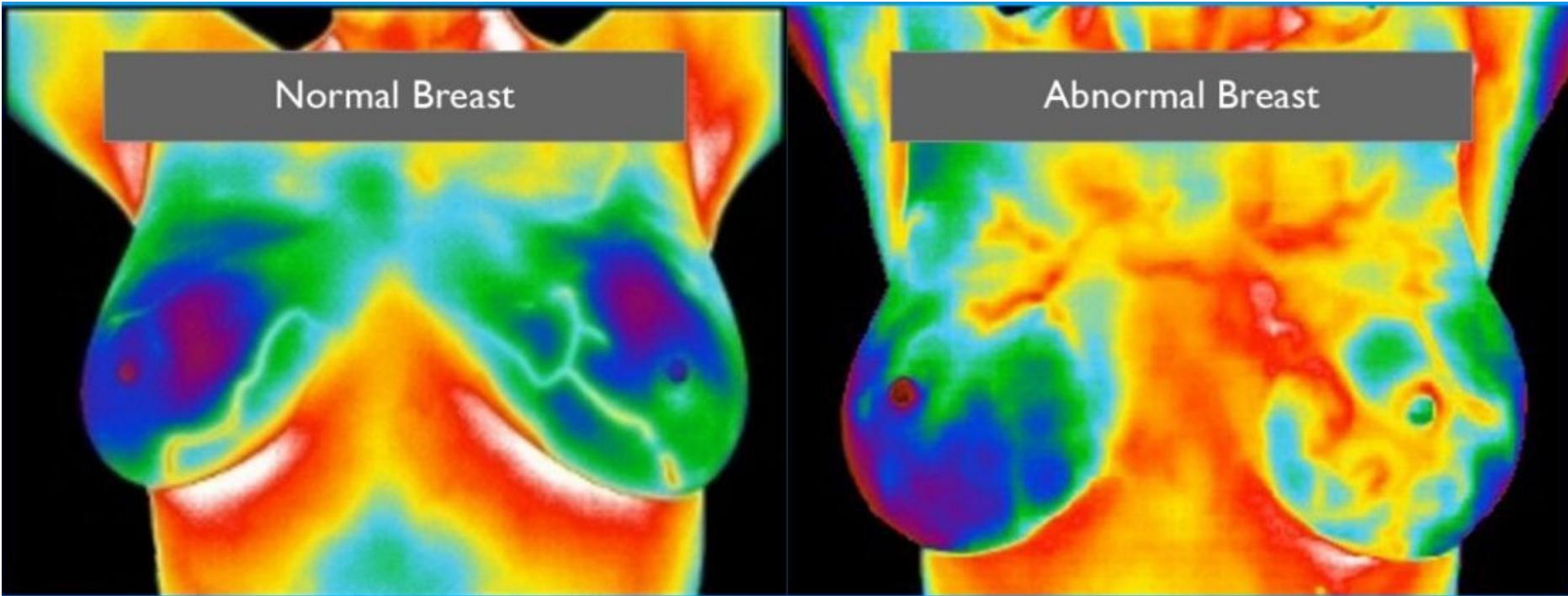


Búsqueda y rescate

# Aplicaciones en Salud

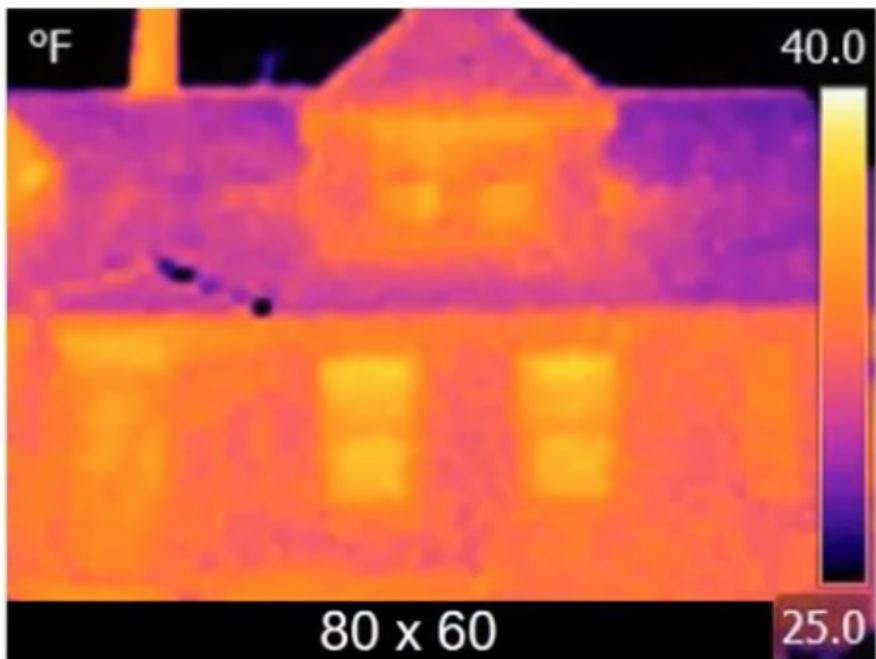
Normal Breast

Abnormal Breast



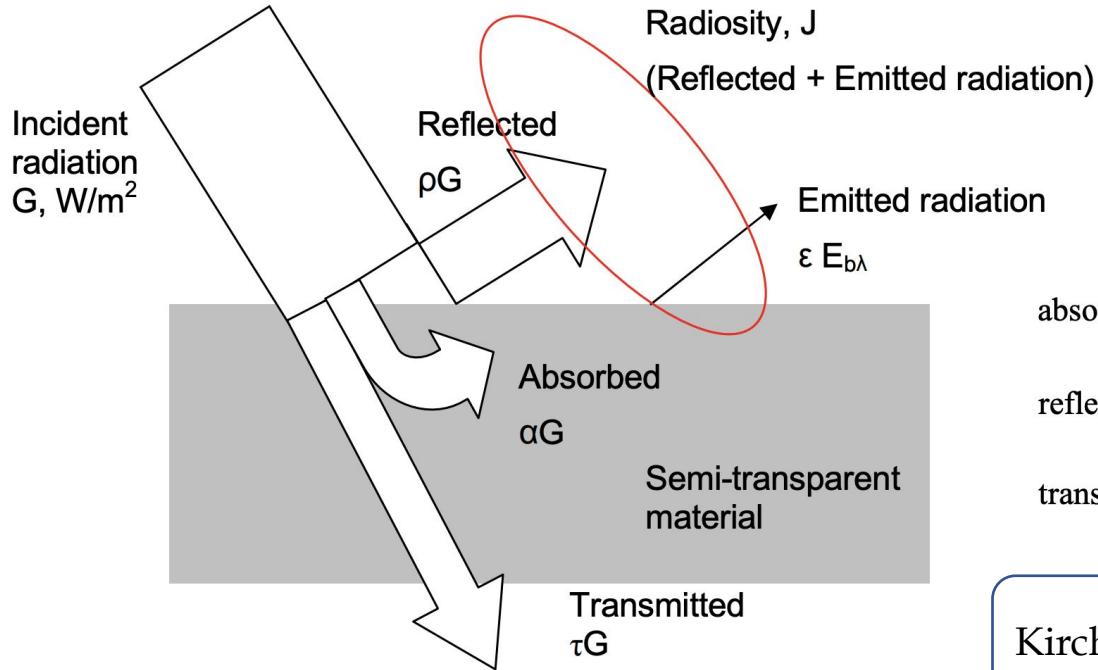


# Efecto de resolución



## 2. Luz y radiación

# Radiation



By first law of thermodynamics  
(conservation of energy):

$$G_{\text{abs}} + G_{\text{ref}} + G_{\text{tr}} = G$$
$$\alpha + \rho + \tau = 1$$

absorptivity :  $\alpha = \frac{\text{absorbed radiation}}{\text{incident radiation}} = \frac{G_{\text{abs}}}{G} \quad 0 \leq \alpha \leq 1$

reflectivity :  $\rho = \frac{\text{reflected radiation}}{\text{incident radiation}} = \frac{G_{\text{ref}}}{G} \quad 0 \leq \rho \leq 1$

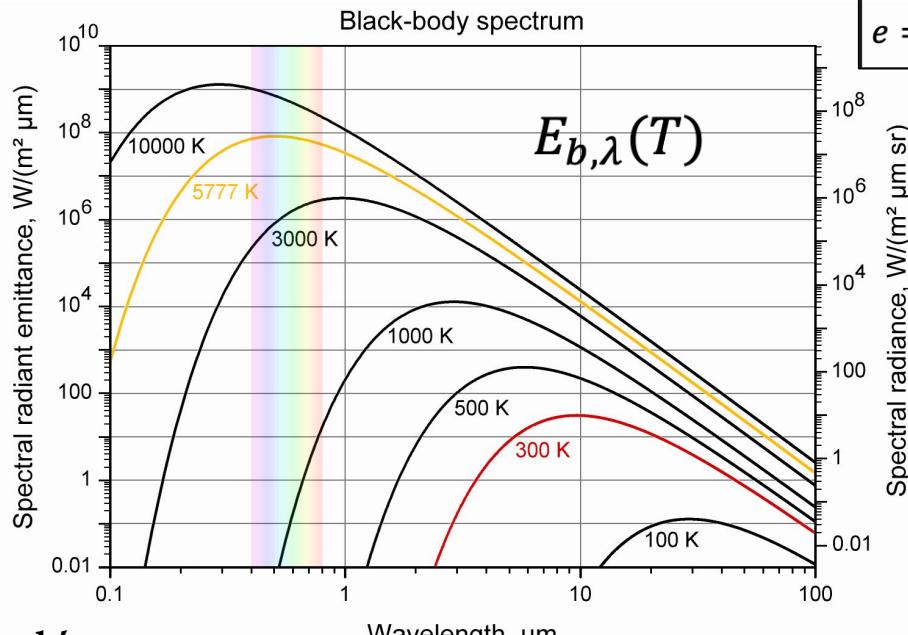
transmissivity :  $\tau = \frac{\text{transmitted radiation}}{\text{incident radiation}} = \frac{G_{\text{tr}}}{G} \quad 0 \leq \tau \leq 1$

Kirchhoff's law:  $\alpha = 1 - \tau, \text{ Transparent } (\rho = 0)$   
 $\alpha = 1 - \rho, \text{ Opaque } (\tau = 0)$

The radiation energy incident on a surface **per unit area** per unit time is called **irradiation**,  $G$ .

# Blackbody radiation

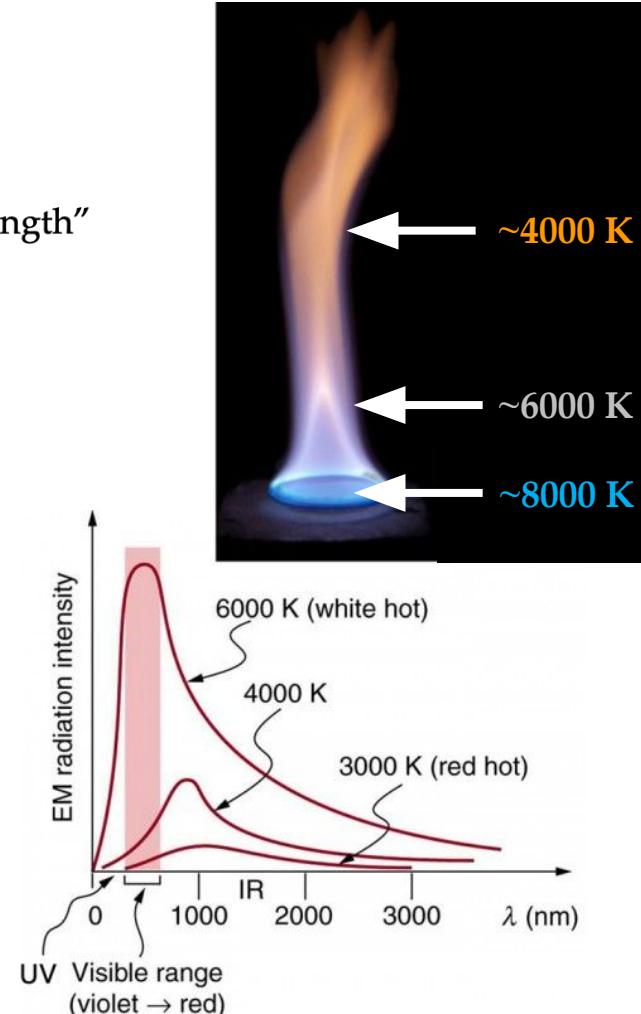
- At a fix  $T$  and  $\lambda$ , no surface can emit more energy than a blackbody
- Einstein: "Energy of a photon is inversely proportional to its wavelength"



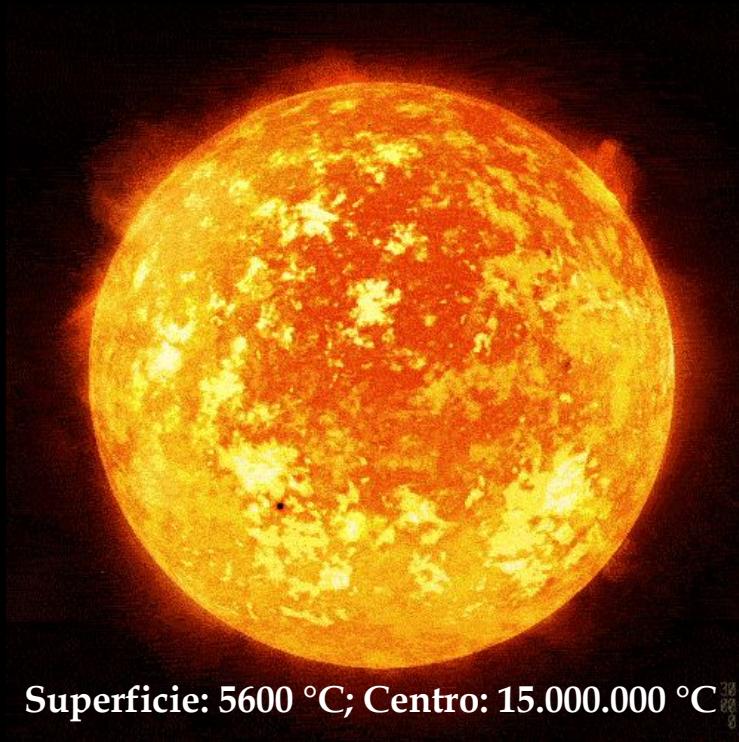
**Planck's law:**

$$E_{b,\lambda}(T) = \frac{2hc^2}{\lambda^5(e^{\frac{hc}{\lambda kT}} - 1)} \left( \frac{W}{m^2 \cdot \mu m} \right)$$

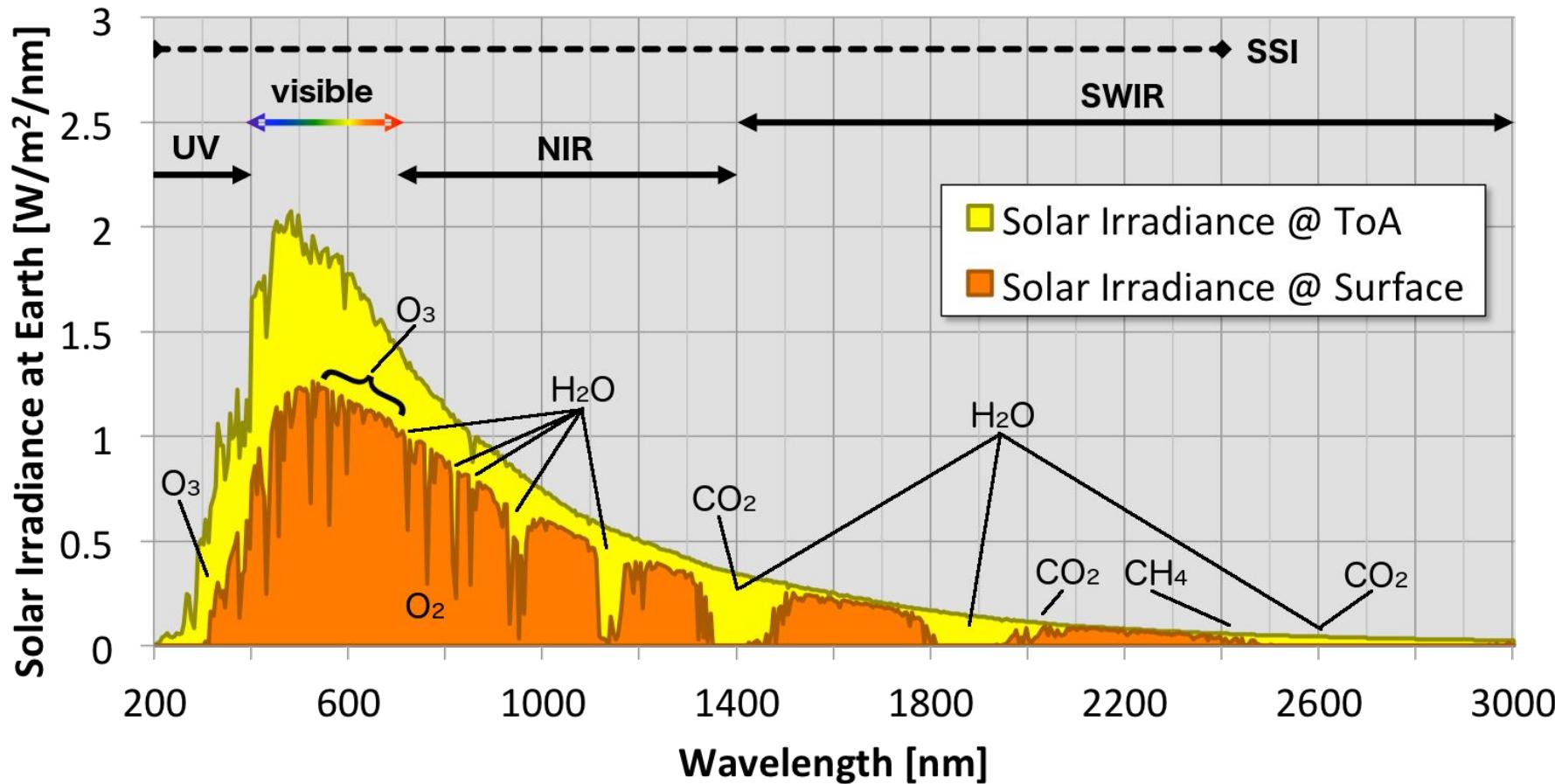
- $c$ : Speed of light
- $h$ : Planck's constant
- $k$ : Boltzmann's constant



# Incandescencia



# Irradiancia Solar



Apparent  
Temperature



Actual  
Temperature



TOTAL

% emission

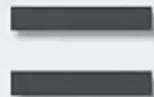
% transmission

% reflection

Reflected  
Apparent  
Temperature

$$\epsilon < 1.0$$

Apparent  
Temperature

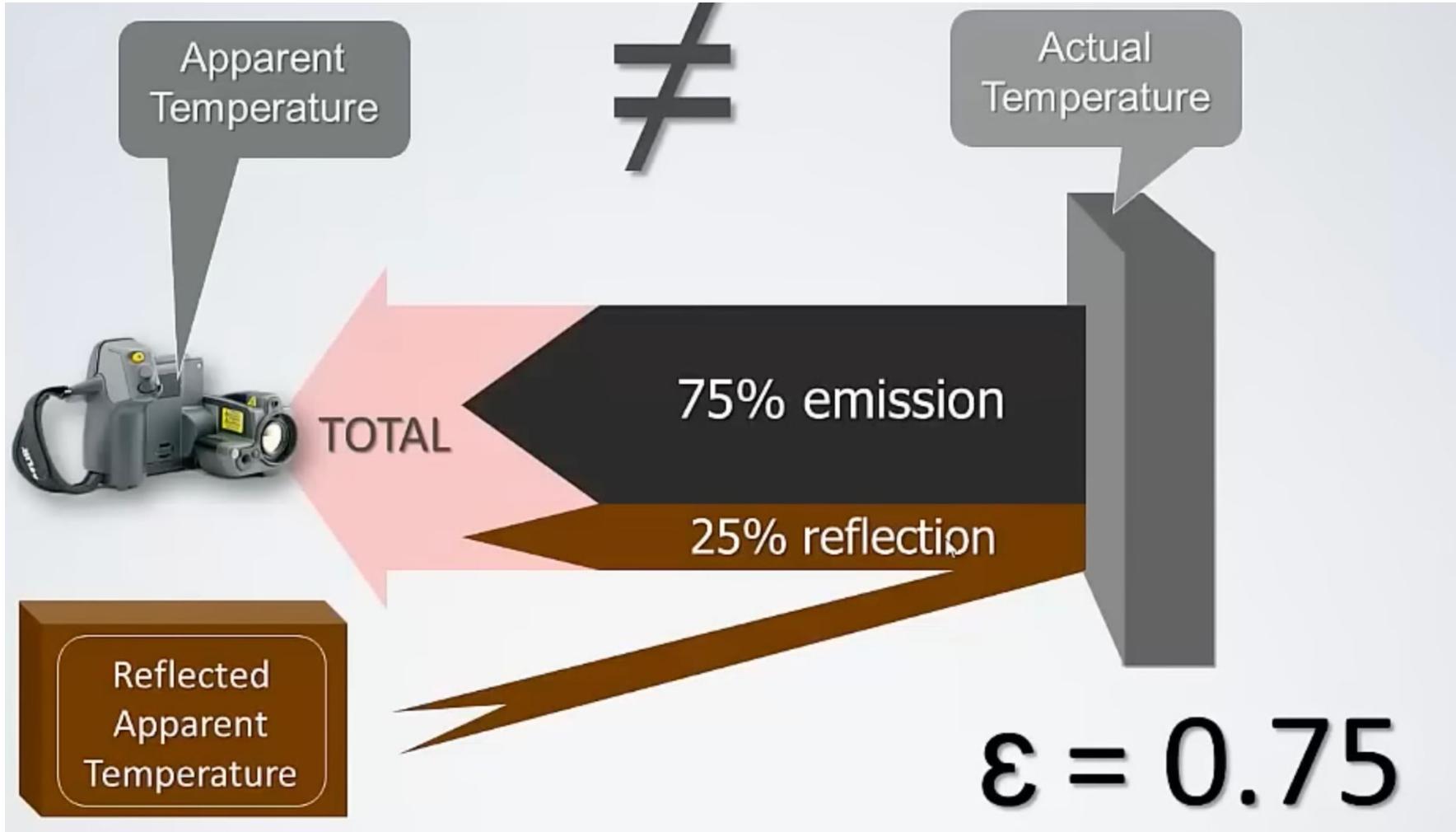


Actual  
Temperature



100% emission

$$\epsilon = 1.0$$



Apparent  
Temperature

Actual  
Temperature  
40°C



TOTAL

75% emission

25% reflection



$$\epsilon = 0.75$$

Apparent  
Temperature

>

Actual  
Temperature  
40°C



TOTAL

75% emission

25% reflection



$$\epsilon = 0.75$$

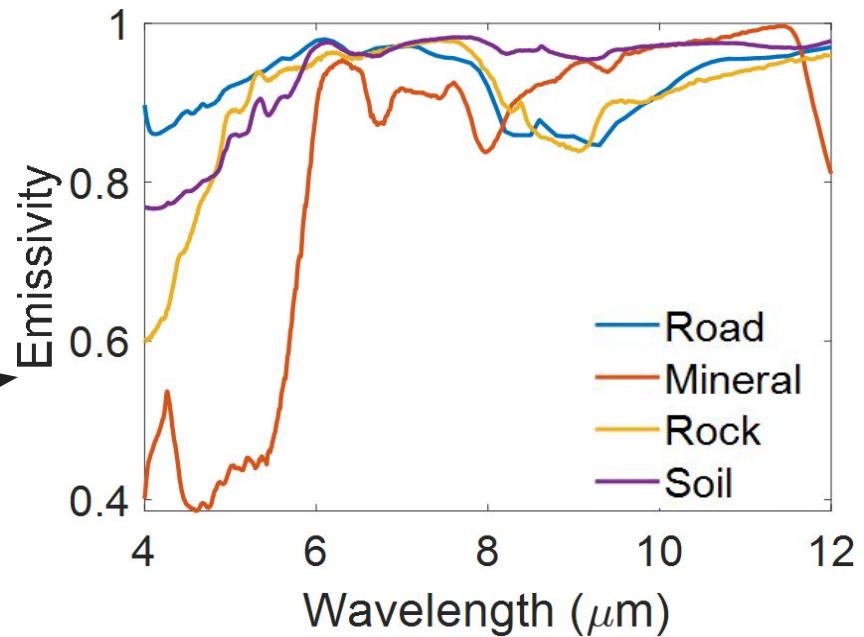
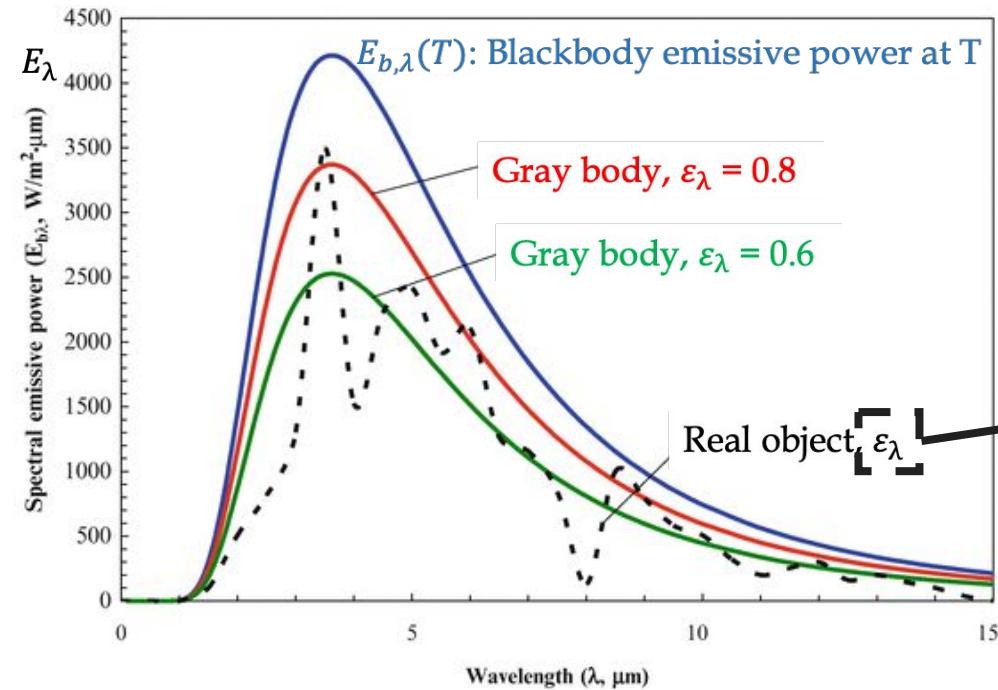
# Emissivity de materiales comunes

Metal	Emissivity	Non-metal	Emissivity
Bare aluminum	0.02–0.4	Concrete (rough)	0.93–0.96
Gold	0.02–0.37	Glass	0.76–0.94
Copper	0.02–0.74	Wood	0.8–0.95
Lead	0.06–0.63	Carbon	0.96
Brass	0.03–0.61	Human skin	0.98
Nickel	0.05–0.46	Paper	0.7–0.95
Steel	0.07–0.85	Plastic	0.8–0.95
Tin	0.04–0.08	Rubber	0.86–0.94
Silver	0.01–0.07	Water	0.67–0.96
Zinc	0.02–0.28	Sand	0.76–0.9

### 3. Imágenes Espectrales Térmicas

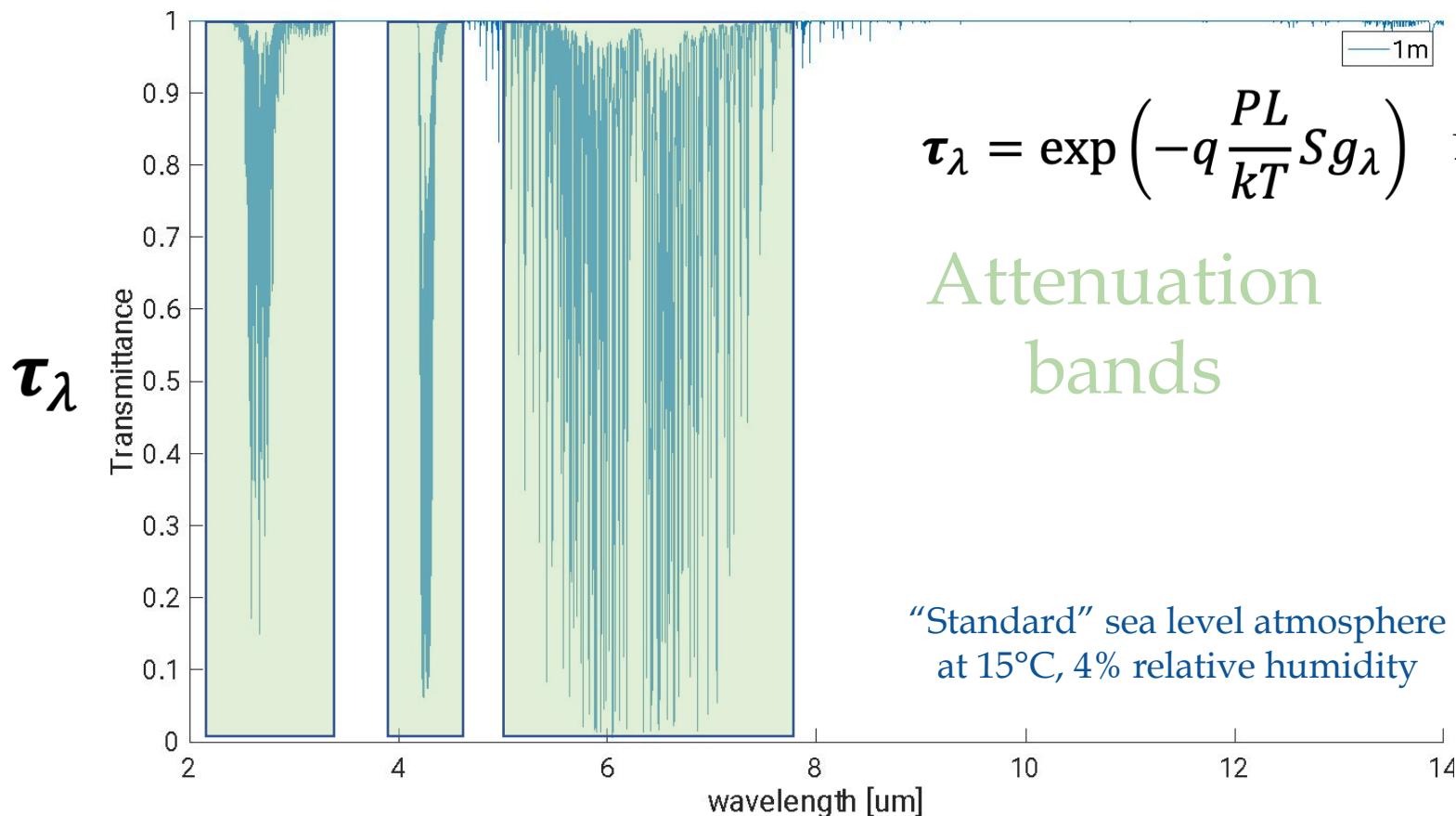
# Emissivity depends on $\lambda$

$$E_\lambda(T) = \varepsilon_\lambda E_{b,\lambda}(T)$$

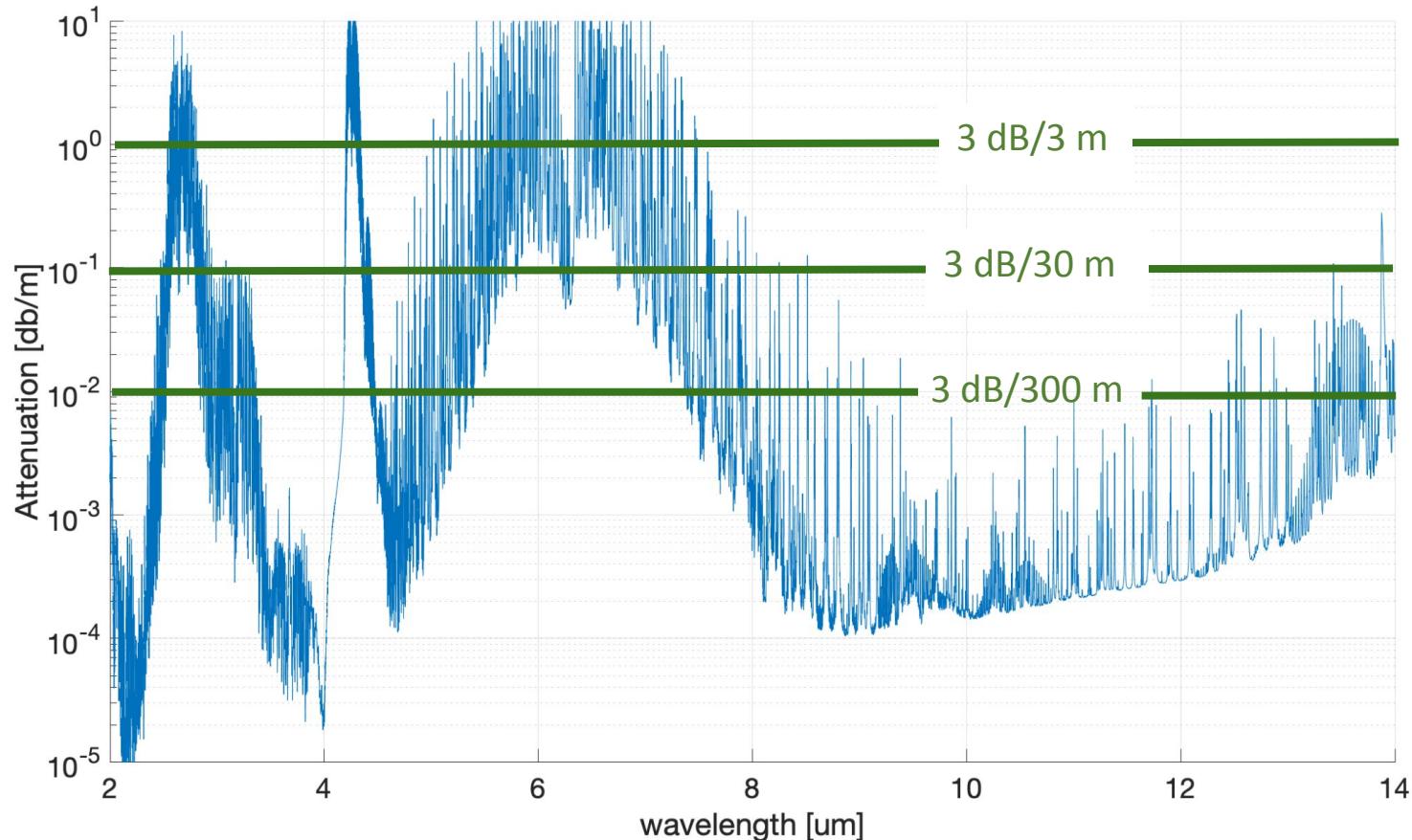




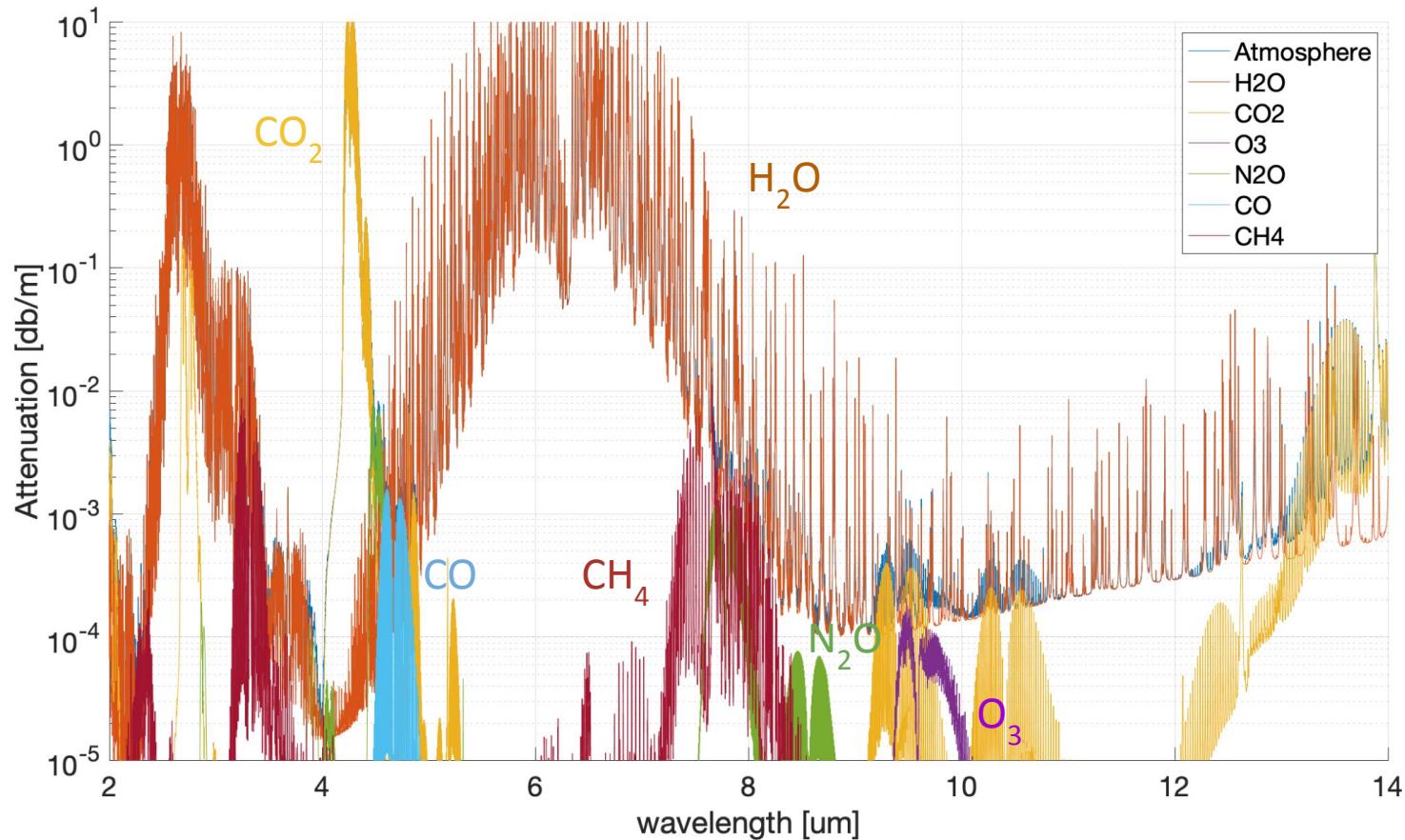
# Atmospheric transmittance



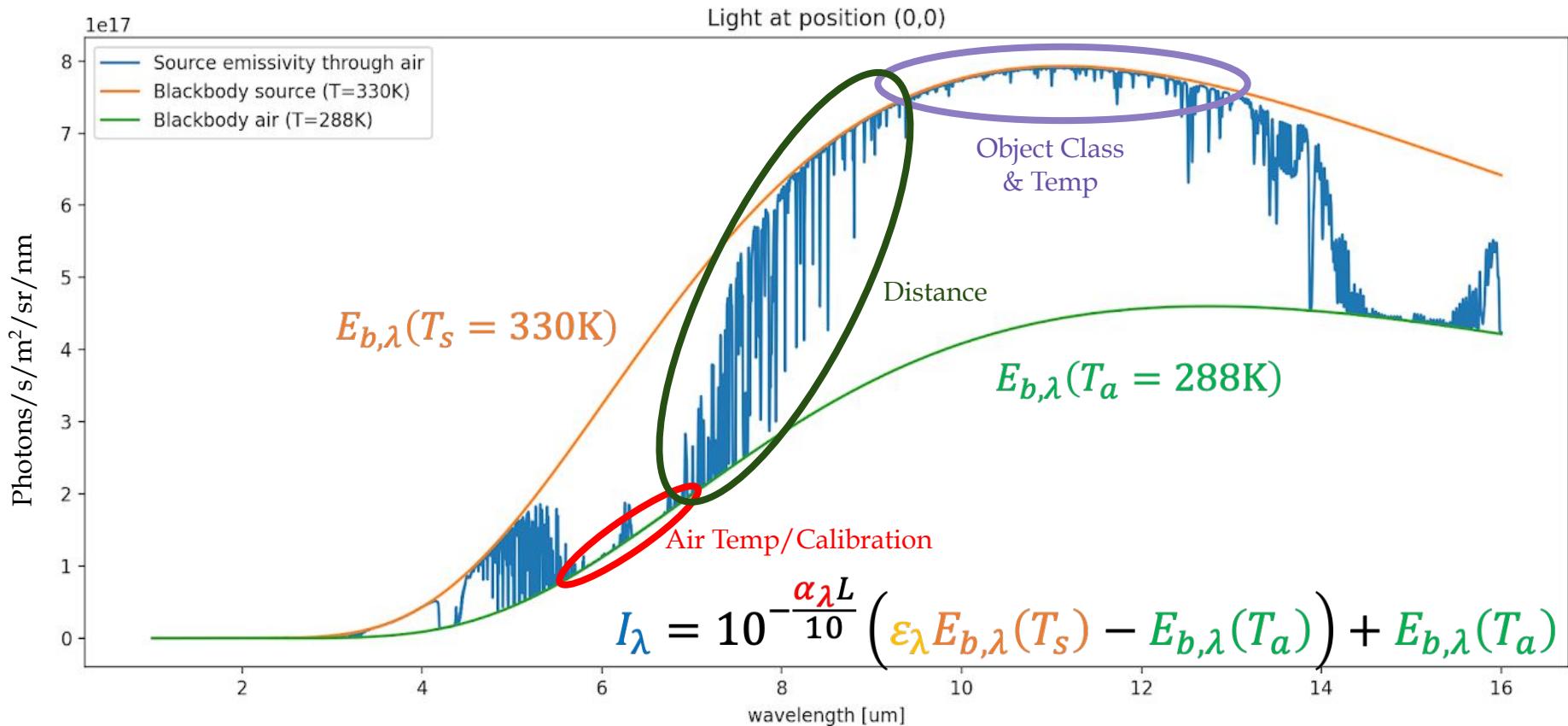
# Atmospheric attenuation



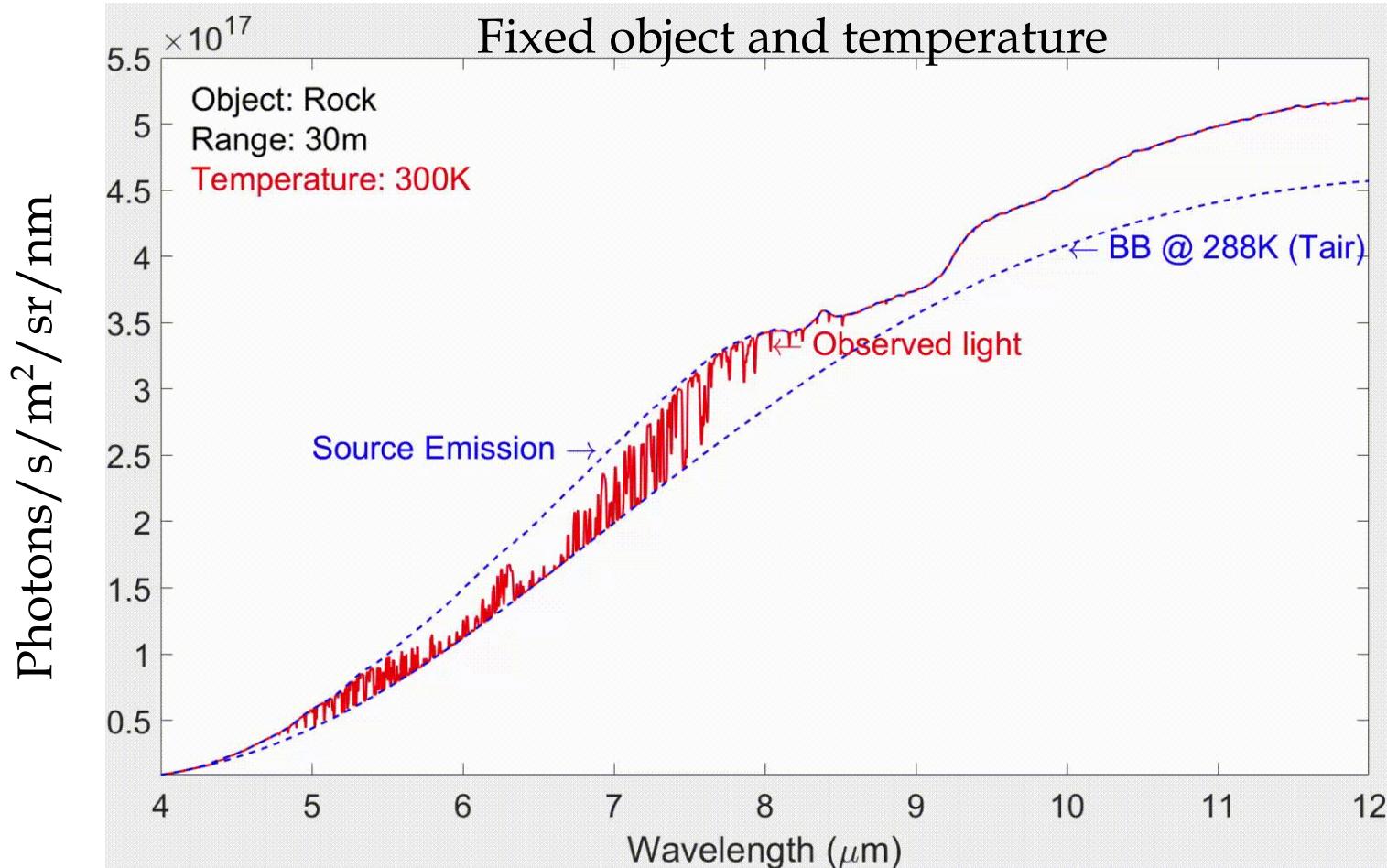
# Atmospheric attenuation by gas composition



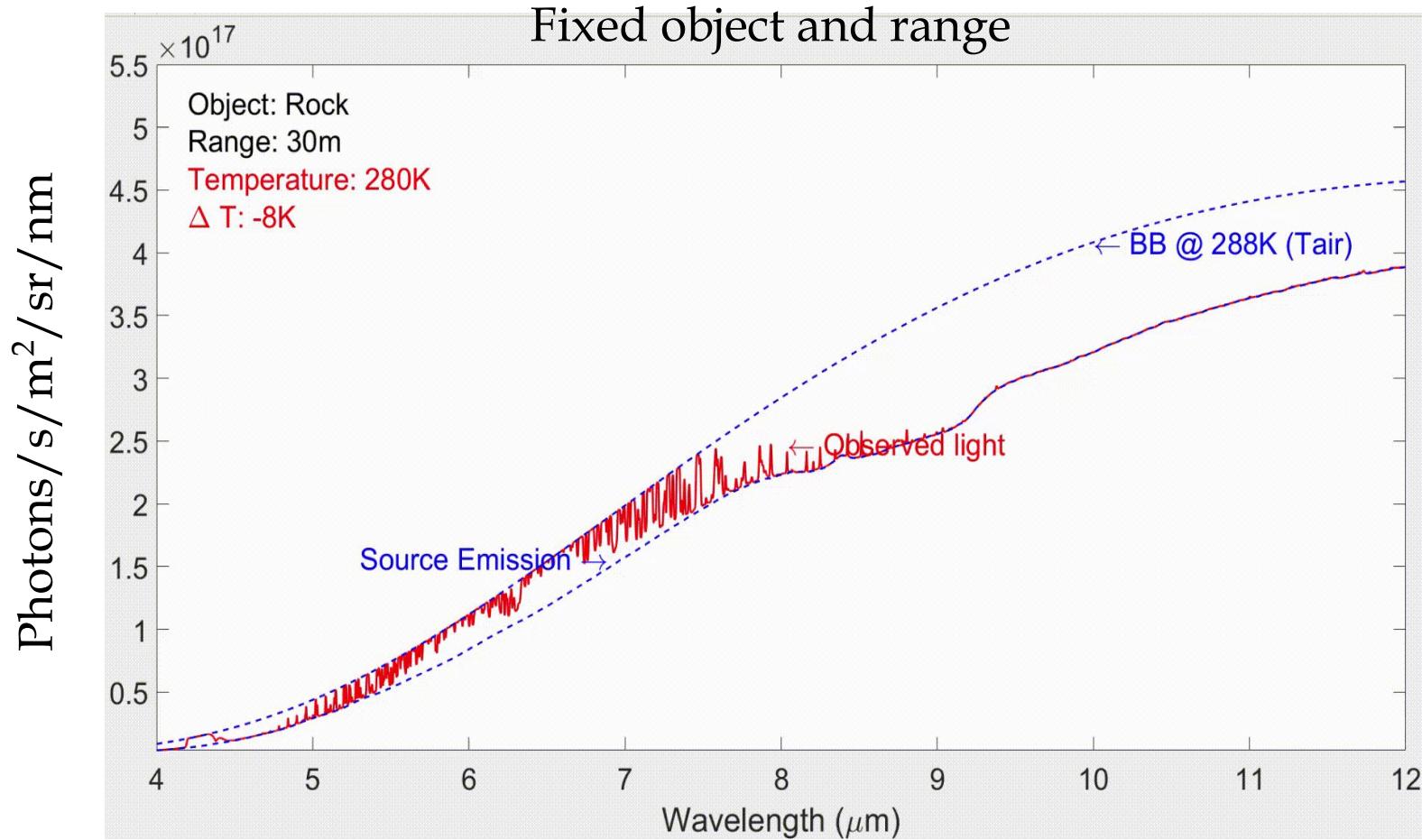
# Emissivity through air



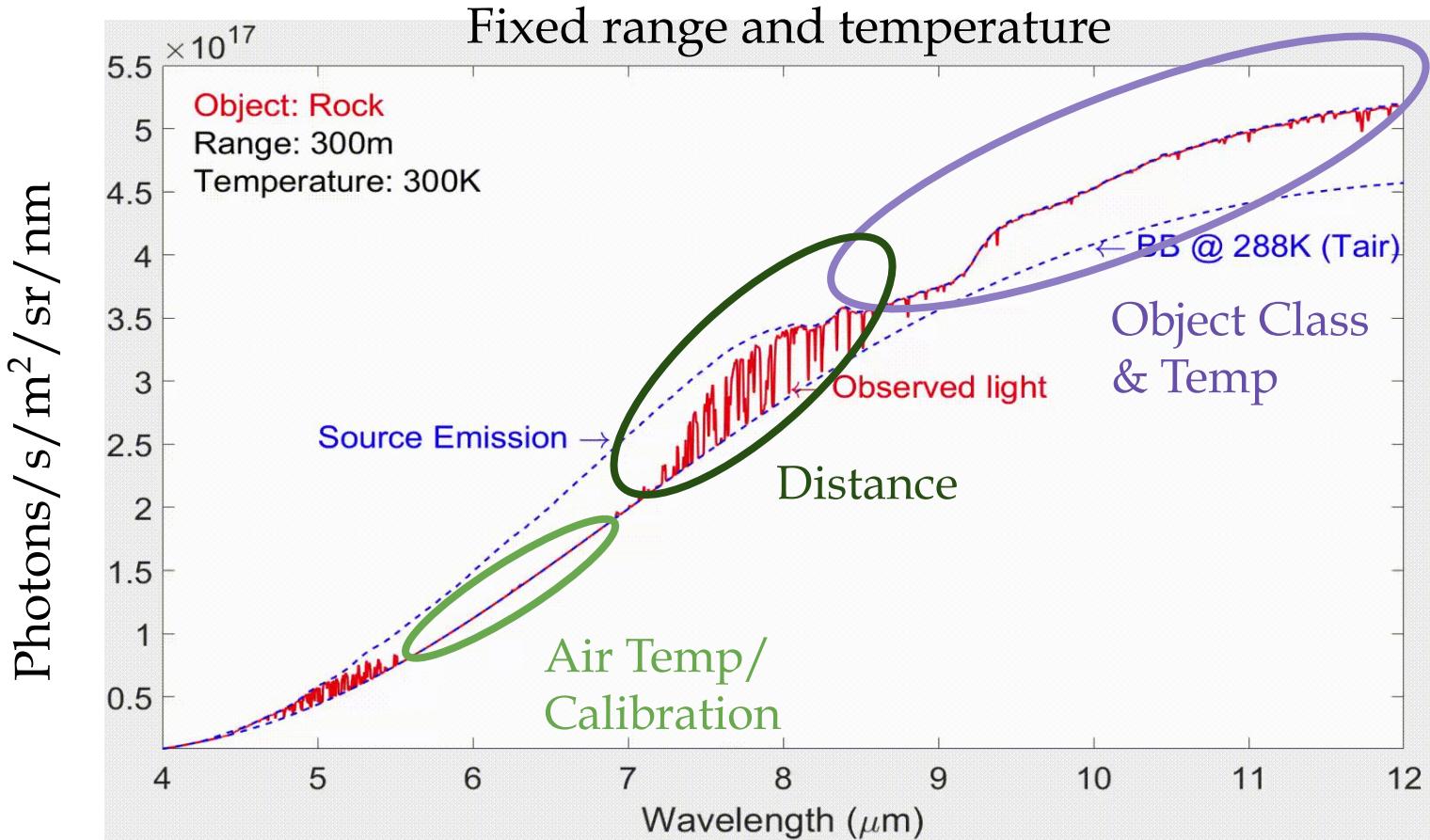
# Range variation



# Temperature variation



# Emissivity variation



## 4. Thermal & CV

<https://www.nature.com/articles/s41586-023-06174-6>

The international journal of science / 27 July 2023



### Turing point

The hunt for better ways to test the intelligence of chatbots

### A place out of the sun

Why planning for shade is key to keeping cool in hot cities

### Parasite restrained

A sustainable way to tackle the snails behind schistosomiasis

Net ref. No. 7071  
Eikon@nature.com

# TeX Vision: Temperature, emissivity, teXture

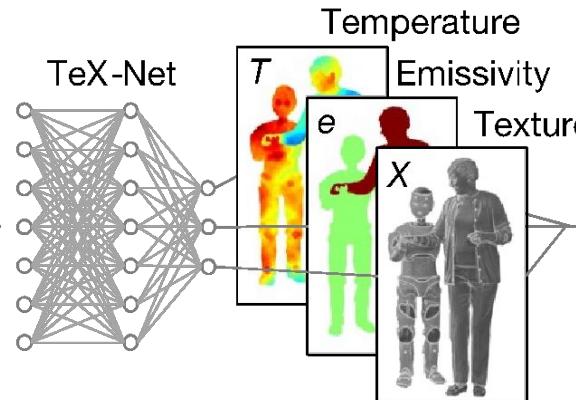
Ghosting effect  
(TeX degeneracy)



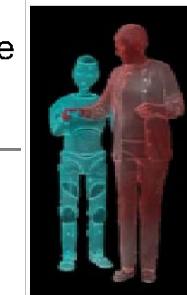
Hyperspectral  
heat cube



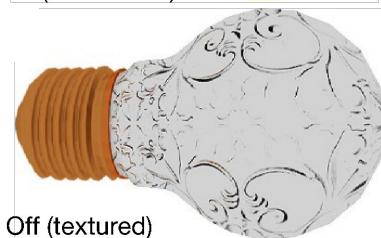
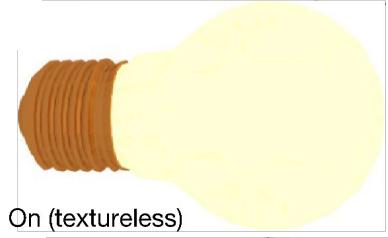
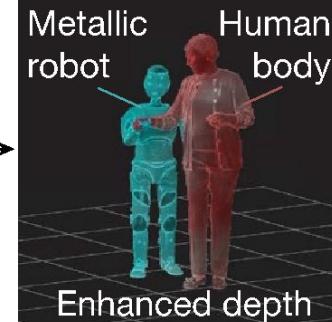
HADAR



TeX vision



HADAR perception



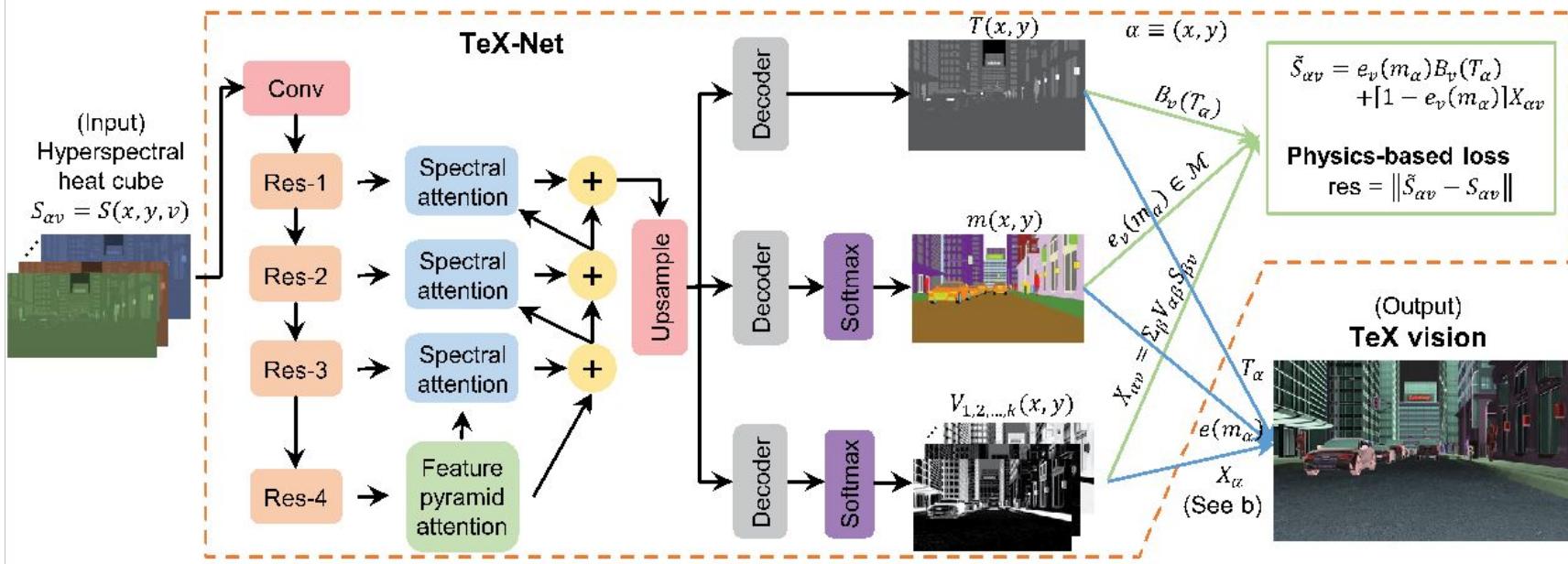
$$I_{\lambda} = 10^{-\frac{\alpha_{\lambda} L}{10}} \left( \varepsilon_{\lambda} E_{b,\lambda}(T_s) - E_{b,\lambda}(T_a) \right) + E_{b,\lambda}(T_a)$$

$$I_{\lambda} = 10^{-\frac{\alpha_{\lambda} L}{10}} \left( \varepsilon_{\lambda} E_{b,\lambda}(T_s) \right) + \left( 1 - 10^{-\frac{\alpha_{\lambda} L}{10}} \right) E_{b,\lambda}(T_a)$$

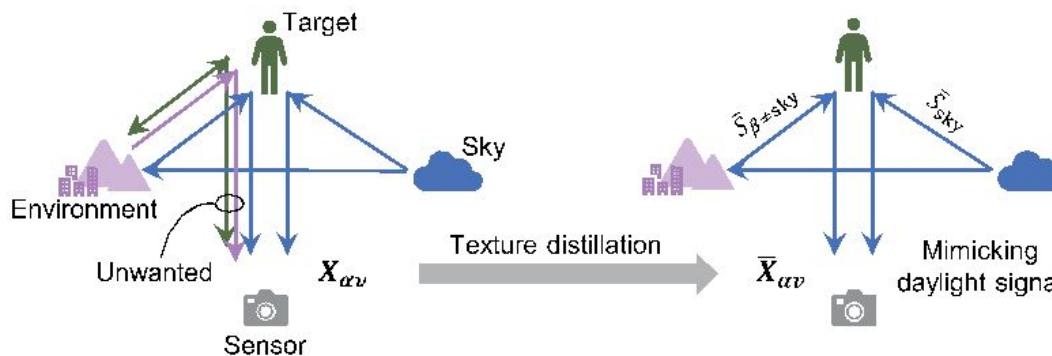
$$S_{\alpha\nu} = e_{\alpha\nu} B_{\nu}(T_{\alpha}) + [1 - e_{\alpha\nu}] X_{\alpha\nu}$$

# T e X - N e t

**a Inverse decomposition**



**b Forward rendering**

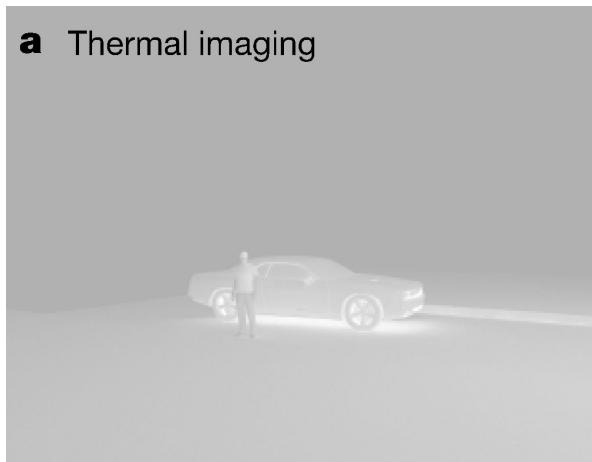


Forward texture distillation procedures:

- (1)  $\bar{S}_{\gamma v} \equiv \begin{cases} S_{\beta v} & \beta = \gamma, \\ 0 & \text{otherwise} \end{cases} \quad \gamma = \text{sky}$
- (2)  $\bar{S}_{\beta v} \equiv [1 - e_v(m_\beta)] \cdot \sum_\gamma V_{\beta \gamma} \bar{S}_{\gamma v}, \quad \beta \neq \text{sky}$
- (3)  $\bar{X}_{\alpha v} \equiv \sum_\beta V_{\alpha \beta} \bar{S}_{\beta v}$
- (4)  $X_\alpha \equiv \sum_v \bar{X}_{\alpha v} \oplus res$

# HADAR: Fundamental limits in ranging

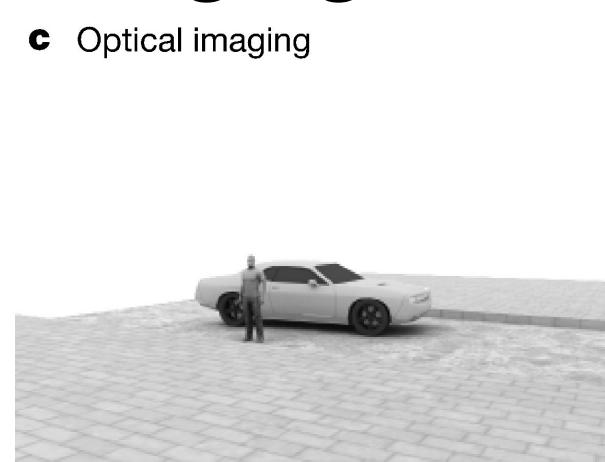
**a** Thermal imaging



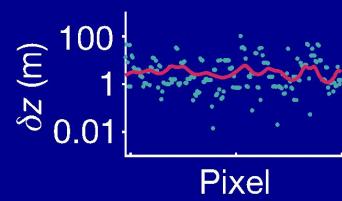
**b** HADAR



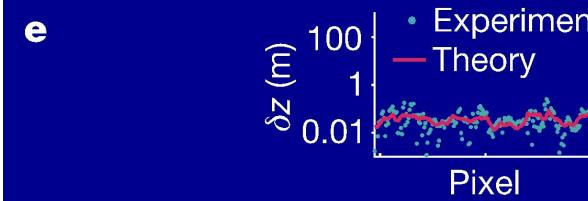
**c** Optical imaging



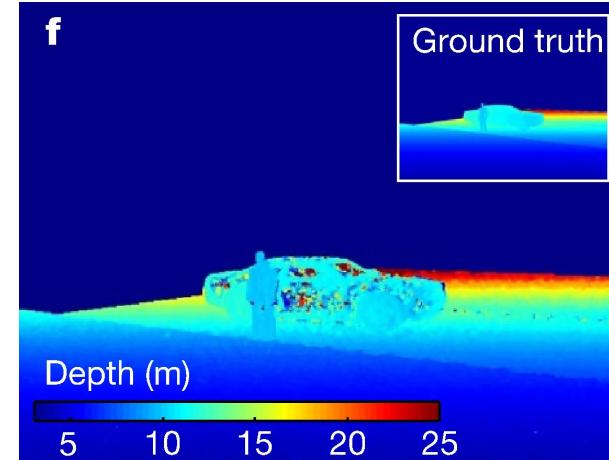
**d**



**e**

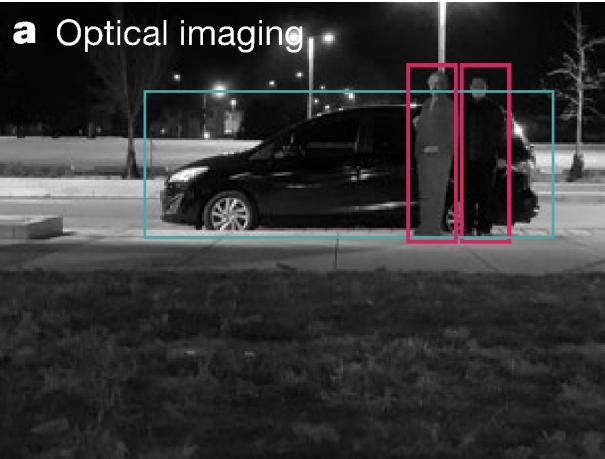


**f**

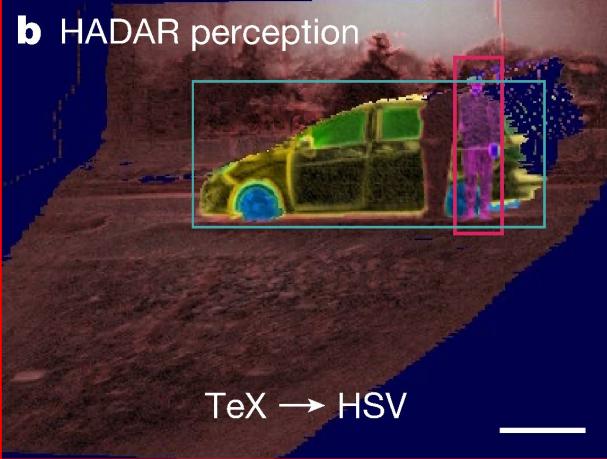


# HADAR: Physics-driven perception

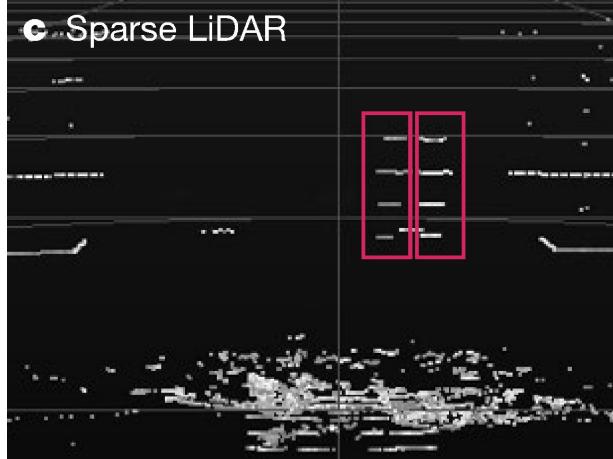
**a** Optical imaging



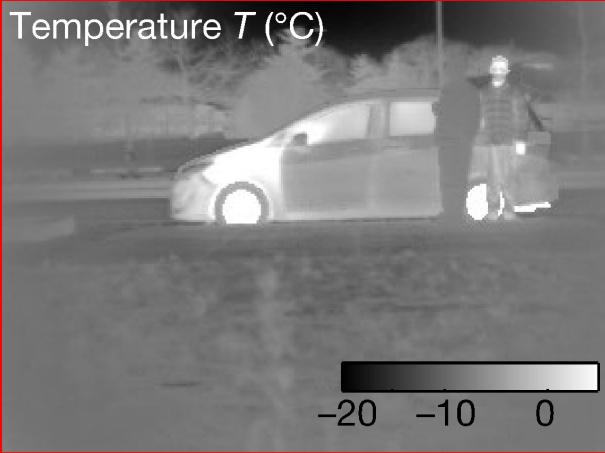
**b** HADAR perception



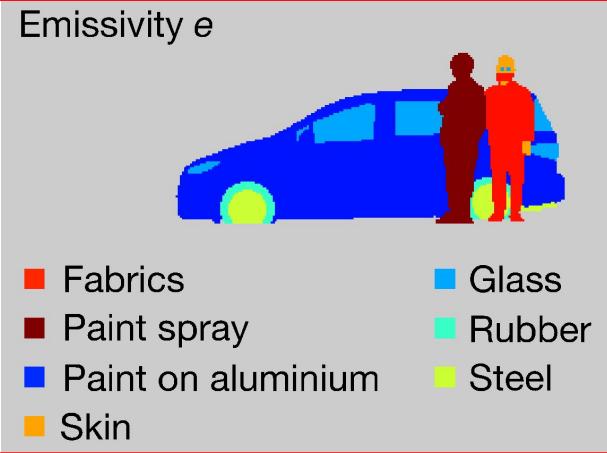
**c** Sparse LiDAR



Temperature  $T$  ( $^{\circ}\text{C}$ )



Emissivity  $e$



Texture X

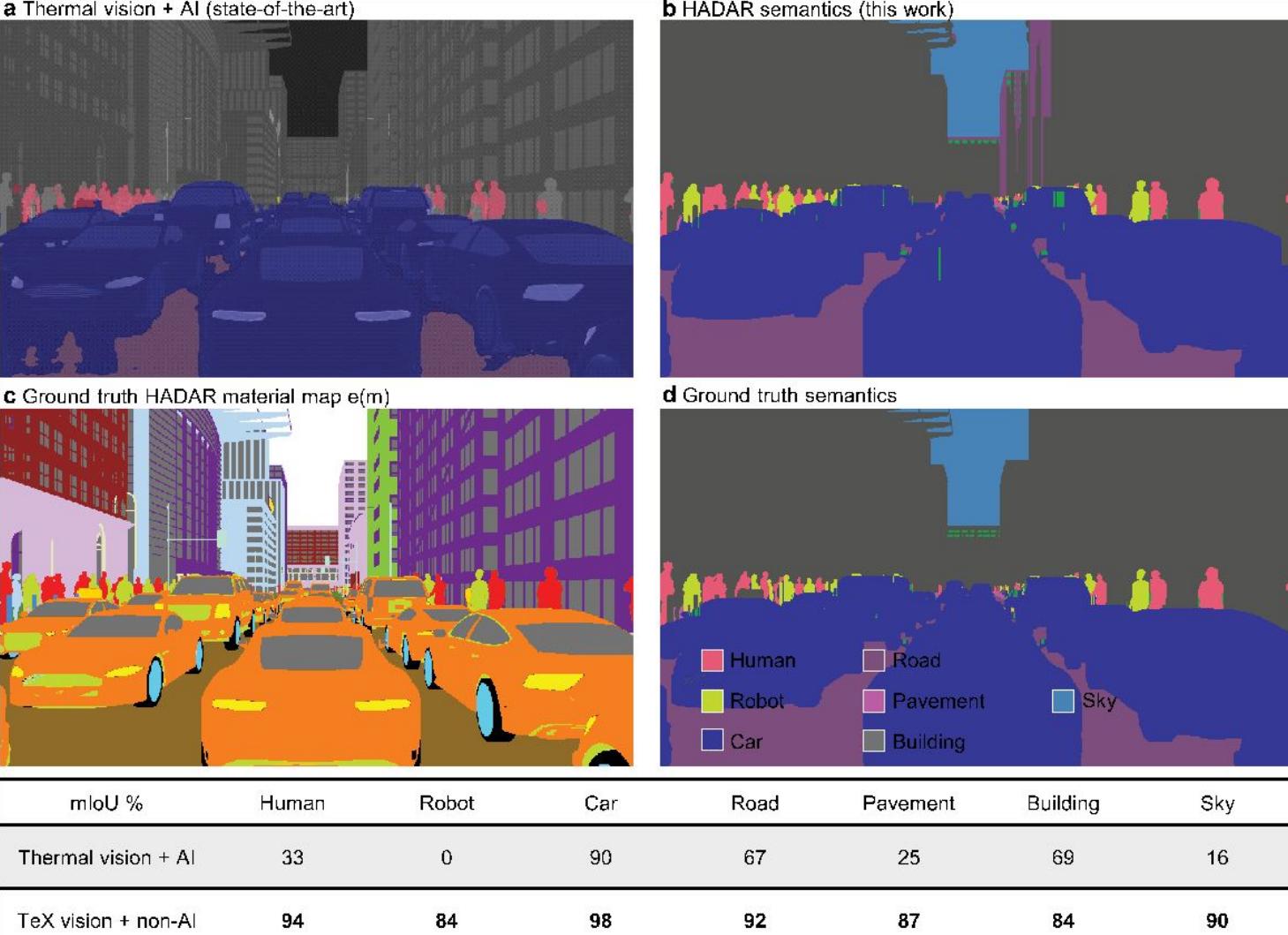


# Thermal Semantic Segment.

(DANet)

Fu, J. et al. in *CVPR*  
3146–3154, 2019.

**mIoU:**  
pixelwise mean  
intersection  
over union.



09-25-2022 Sun 05:43:22



Grill Temp: 451° F

Tri Tip 1 Sausage 1  
Start: 05:40 PM Start: 05:40 PM  
Duration: 03m:55s Duration: 01m:30s  
Temp: 261.4°F Temp: 253.0°F

Tri Tip 2 Sausage 2  
Start: 05:40 PM Start: 05:40 PM  
Duration: 03m:55s Duration: 01m:30s  
Temp: 262.0°F Temp: 254.1°F

Tri Tip 3 Sausage 3  
Start: 05:40 PM Start: 05:40 PM  
Duration: 03m:47s Duration: 01m:30s  
Temp: 261.5°F Temp: 253.0°F

Sausage 4  
Start: 05:40 PM Start: 05:40 PM  
Duration: 03m:38s Duration: 01m:30s  
Temp: 260.4°F Temp: 252.0°F

Cancer 0.1

# Navigation





vehicle

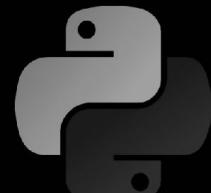
vehicle vehicle vehi vehicle

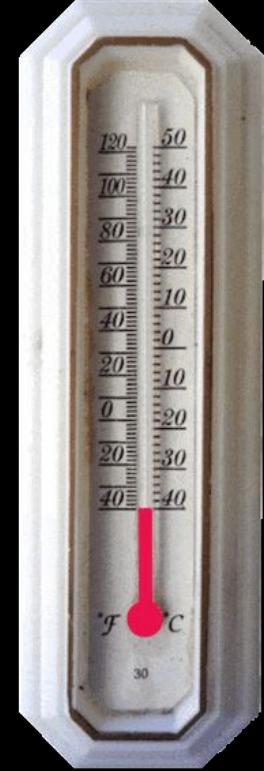
ve vehicle v vehicle cle ; vehicle

vehicle

person person

# 5. Hands-on: Thermal Imaging

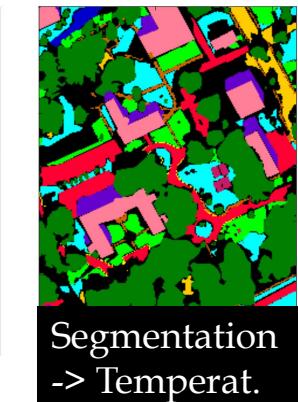
 + colab  
python



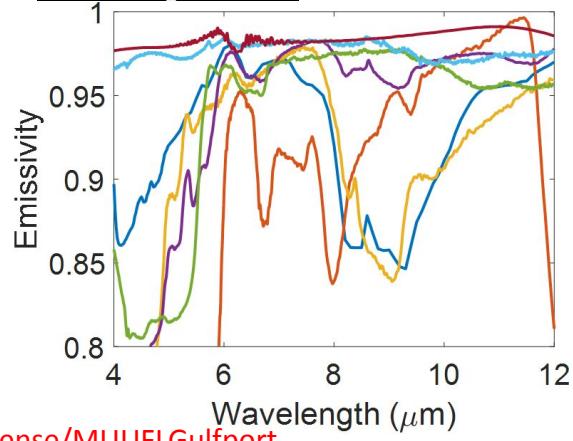
# Computational image formation and analysis



RGB



## Segmentation -> Temperat.



**Toy data after forward model**

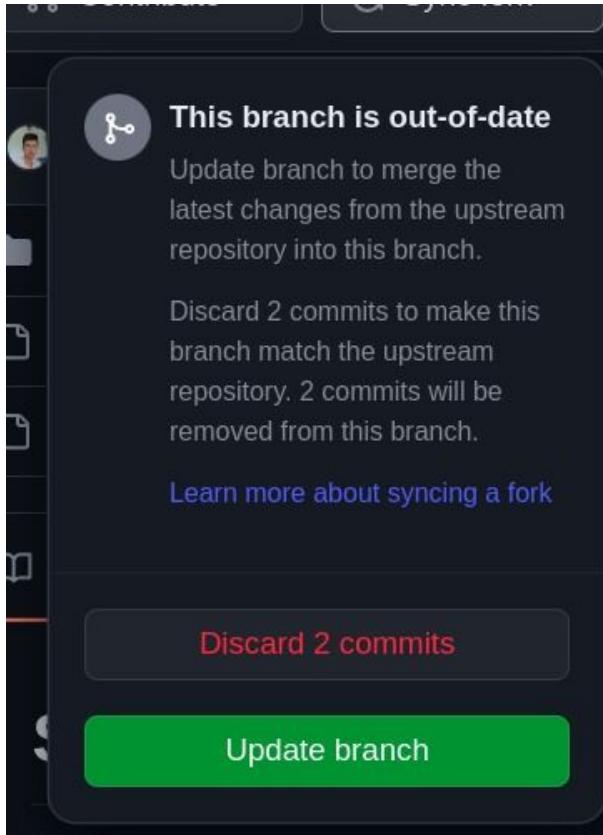
The figure illustrates a forward model for photon counts from a scene. It consists of several components:

- 3D Volume:** A 3D volume rendering of a campus scene. The color of each voxel represents the photon count, as indicated by the color bar on the right.
- 2D Heatmaps:** Two 2D heatmaps showing photon counts at different wavelengths:
  - Photon Count @ 6  $\mu\text{m}$ :** The color bar ranges from 80 to 180.
  - Photon Count @ 8  $\mu\text{m}$ :** The color bar ranges from 100 to 400.
- Line Plot:** A plot of Photon Count versus Wavelength ( $\lambda$ ).
  - The x-axis represents wavelength  $\lambda$  in micrometers ( $\mu\text{m}$ ), ranging from 6 to 12.
  - The y-axis represents the number of photons, ranging from 50 to 200.
  - A sharp peak is visible at approximately 8  $\mu\text{m}$ , labeled "100 kHz".

# Actualicemos el repositorio!

A screenshot of a GitHub repository page. At the top, there are navigation links: 'main' (with a dropdown arrow), '1 Branch' (with a dropdown arrow), '0 Tags' (with a dropdown arrow). To the right are search bar ('Go to file'), a user icon ('t'), 'Add file' (with a dropdown arrow), and a green 'Code' button. Below this, a message states: 'This branch is 2 commits ahead of, 3 commits behind semilleroCV/Hands-on-Computer-Vision:main.' At the bottom left, there are two buttons: 'Contribute' (with a dropdown arrow) and 'Sync fork' (with a dropdown arrow). The 'Sync fork' button is highlighted with a thick red border.

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